

UNIVERSITAT DE BARCELONA



L'Univers proper i llunyà

Estrelles, galàxies i cosmologia

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Universitat de Barcelona –

Institut de Ciències del Cosmos – IEEC

Uc^e

UNIVERSITAT
CATALANA
D'ESTIU

Agost-2007

Estrelles, galàxies i cosmologia

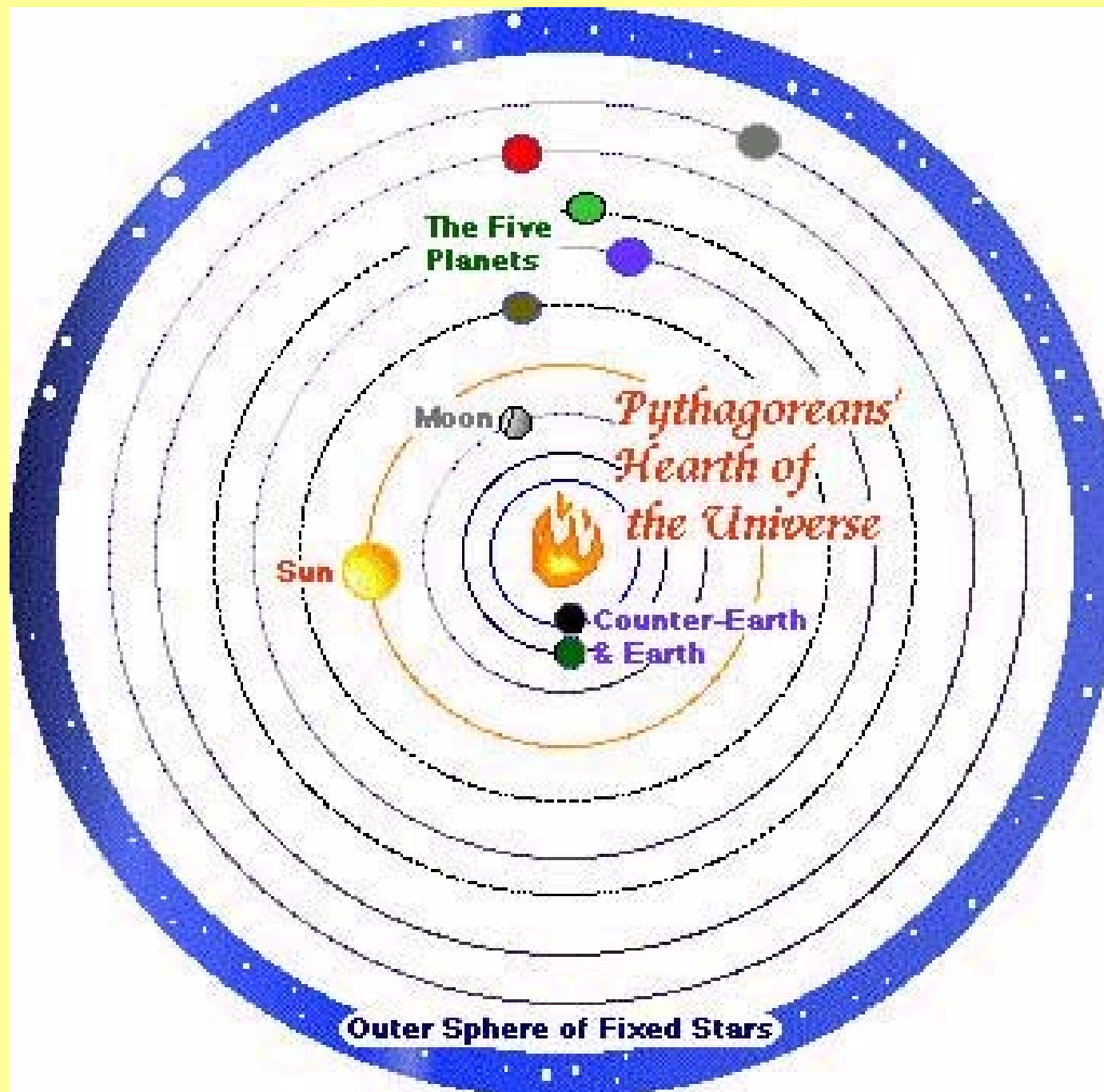
1. Les estrelles

2. El Sol

- L'estrella més propera
- Estudi detallat









Luna

Venus: fases

April 1



May 1



April 10



May 9



April 19

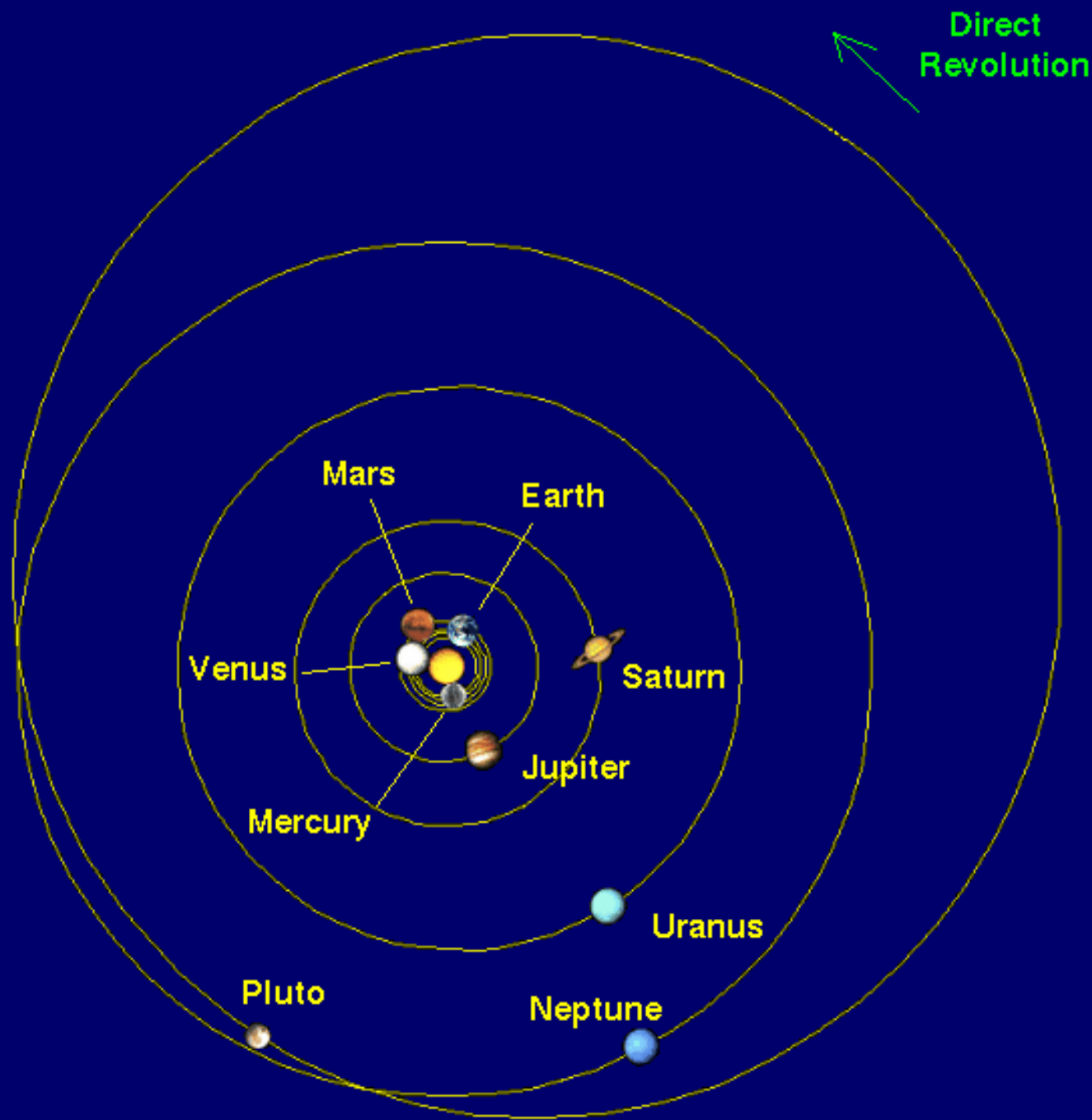


April 25



May 15



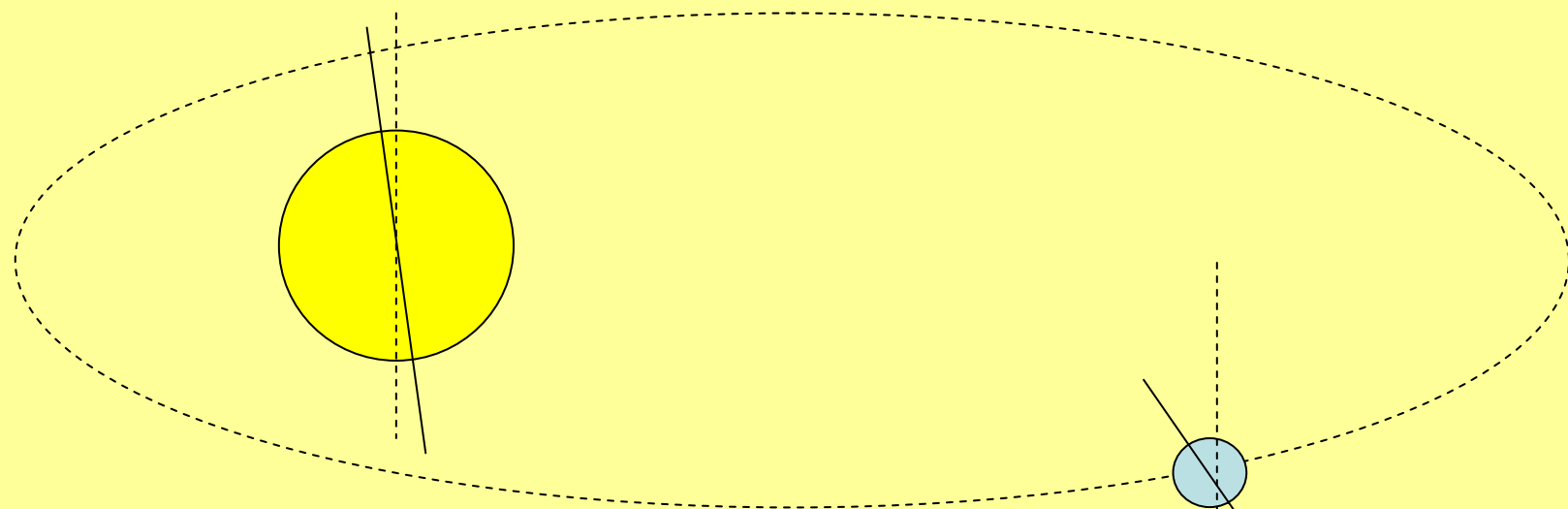


Galileu Galilei

taques

enfosquiment





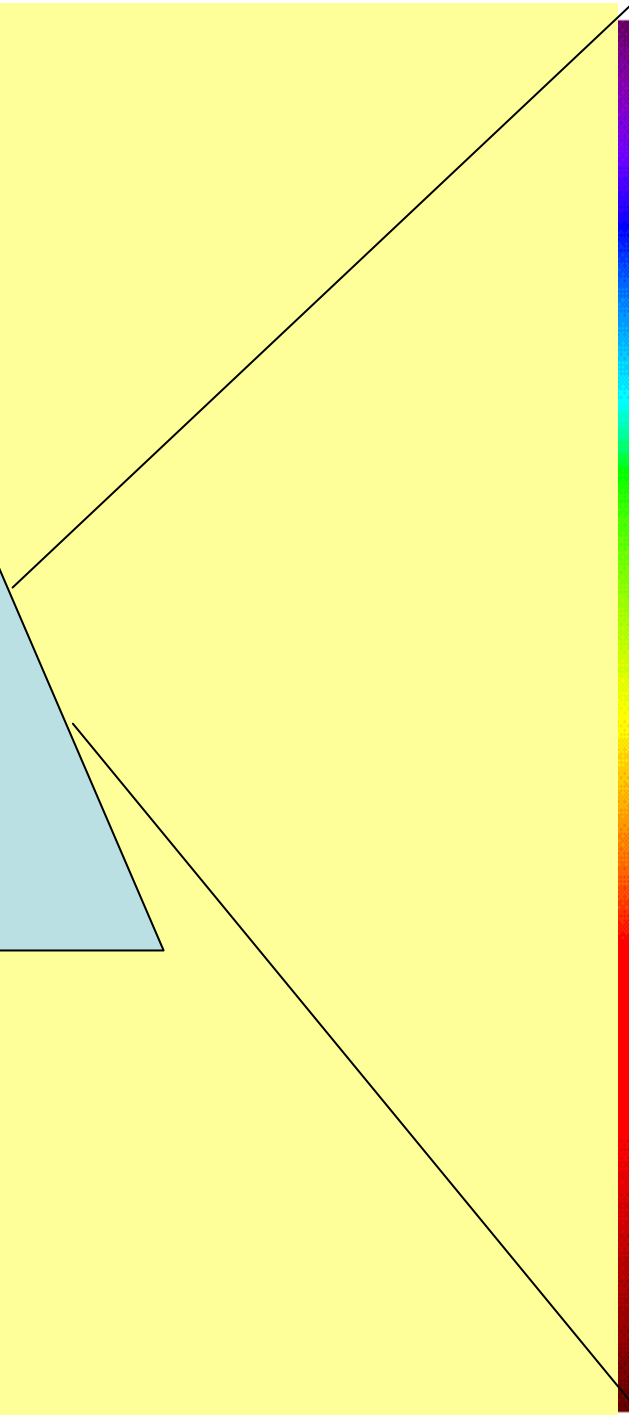
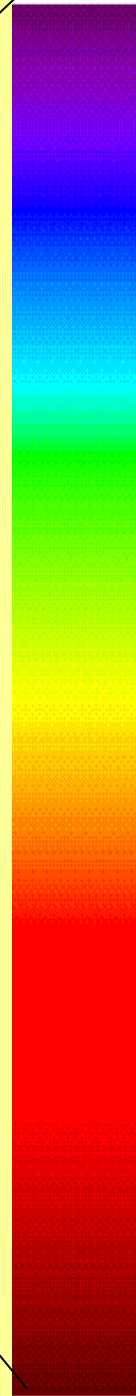
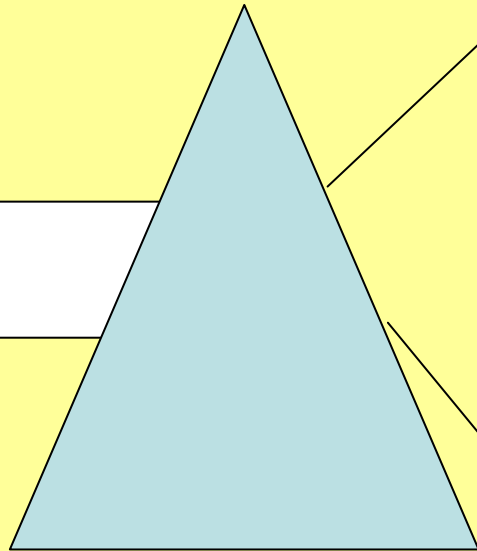
7°

Equador: 25^d
Pols: 34^d

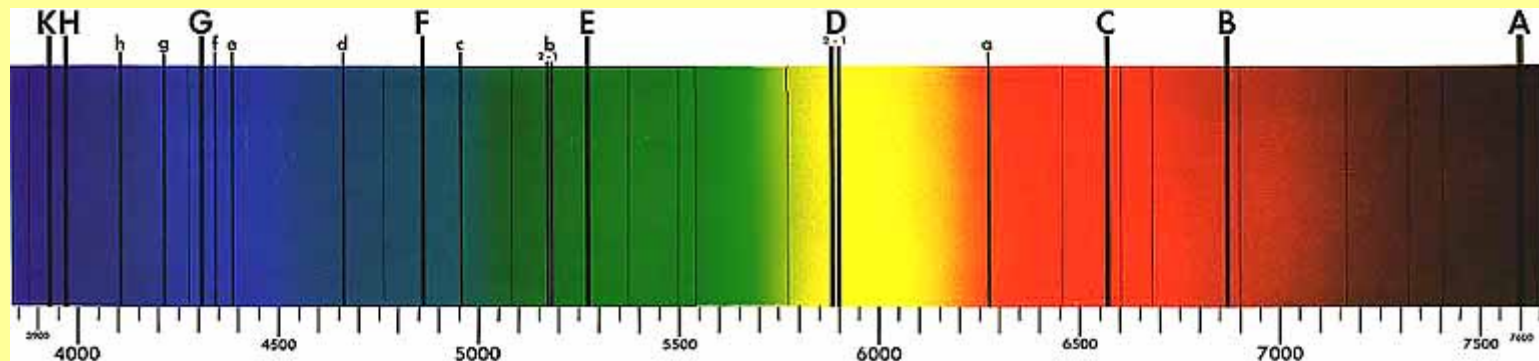
23,5°

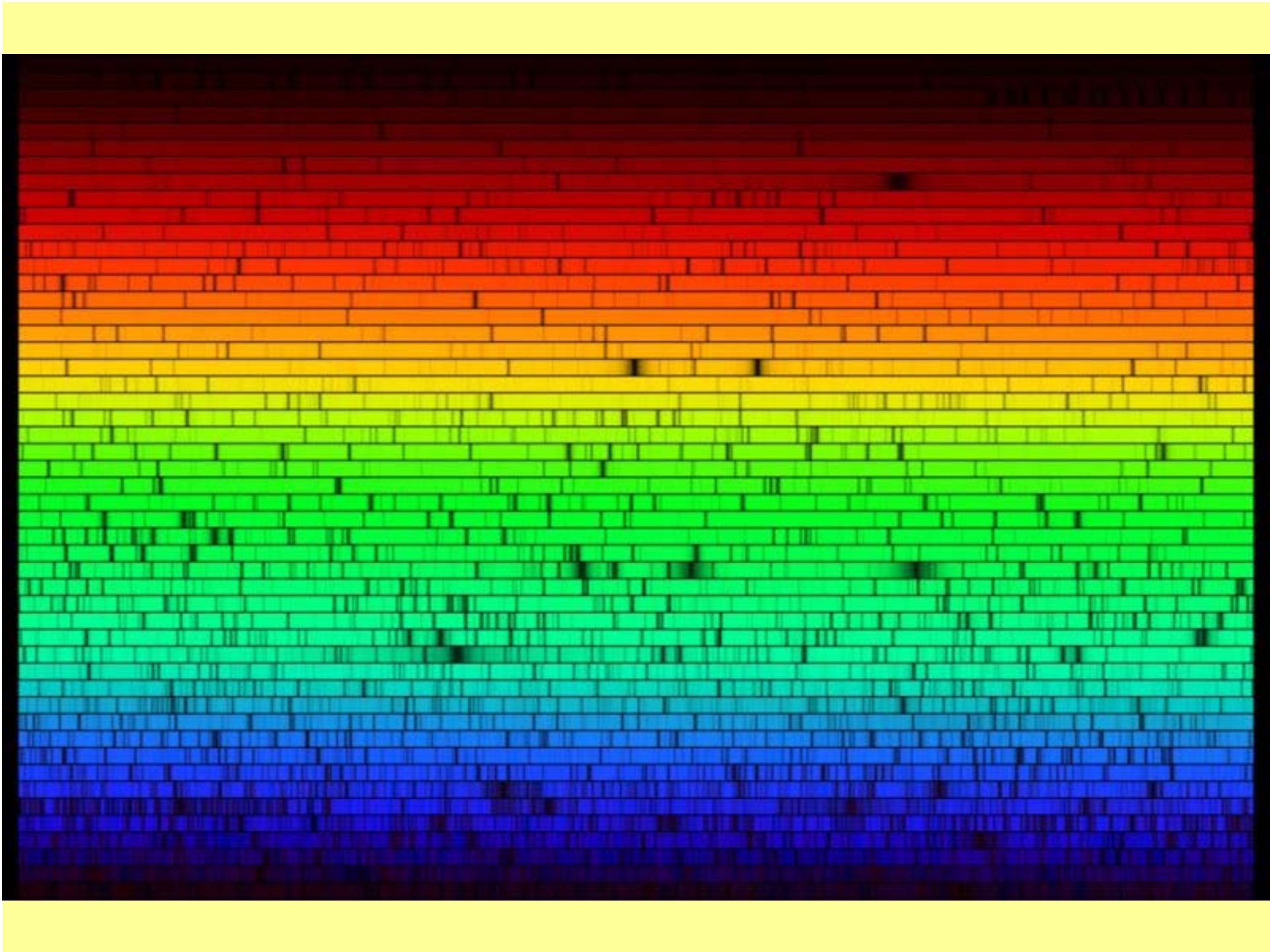
24^h

Newton



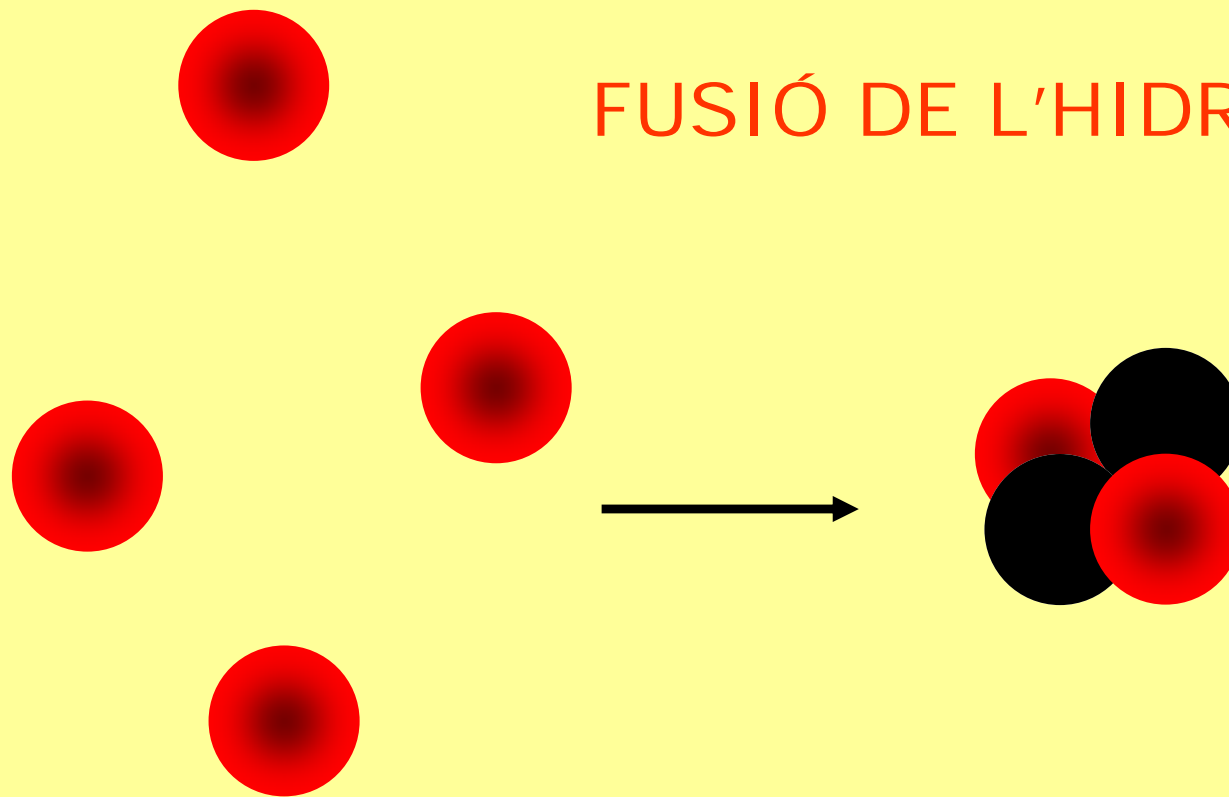
von Fraunhofer





1920: Eddington

FUSIÓ DE L'HIDROGEN



Hidrogen

Heli

i després de tots aquests anys ?

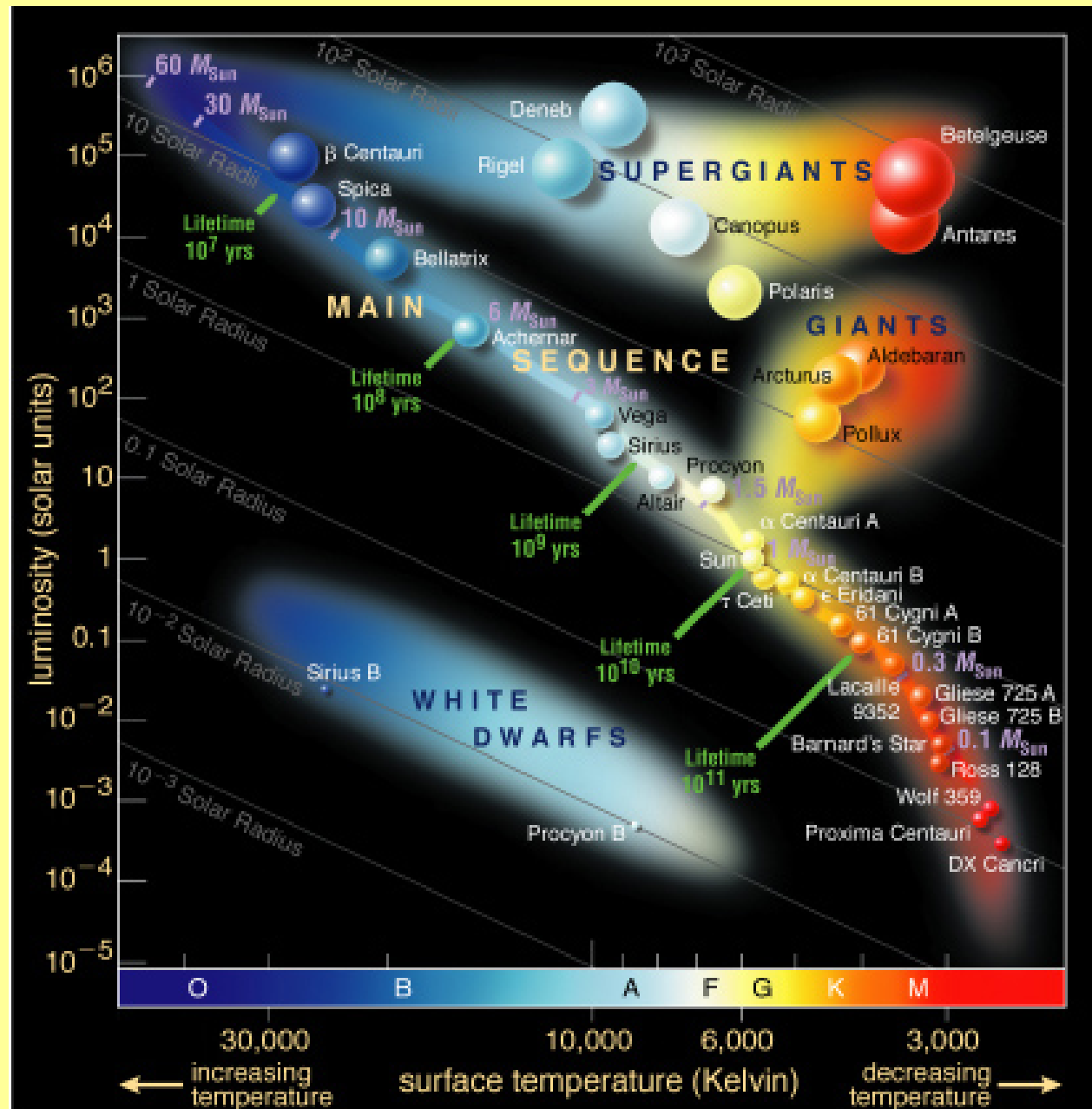
Actualment és a la seqüència principal: fusió hidrogen al nucli

Massa = $2 \cdot 10^{30}$ kg =
333.000 M_{Terra}

Radi (eq) = 700.000 km
= 110 R_{Terra}

Densitat
mitjana = 1400 kg/m^3

gravetat $\sim 28 g_{\text{Terra}}$



Actualment és a la seqüència principal: fusió hidrogen al nucli

Temperatura= 5500 °C

G2 V

Lluminositat= $4 \cdot 10^{26}$ W

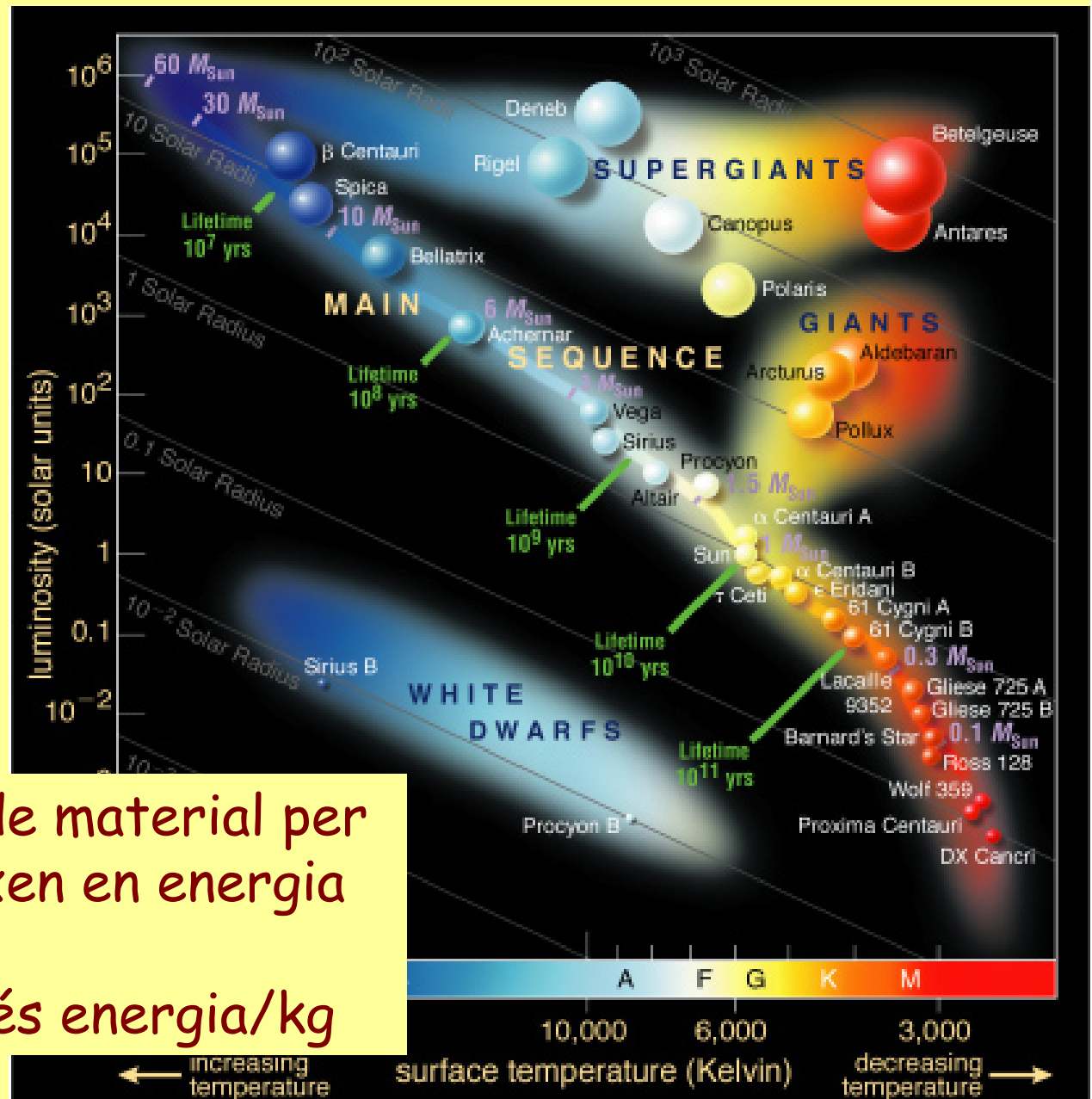
1370 W/m²

V= -26,74 mag

Mv= 4,83

4 milions de tones de material per segon es converteixen en energia

Cos humà: 1 milió més energia/kg

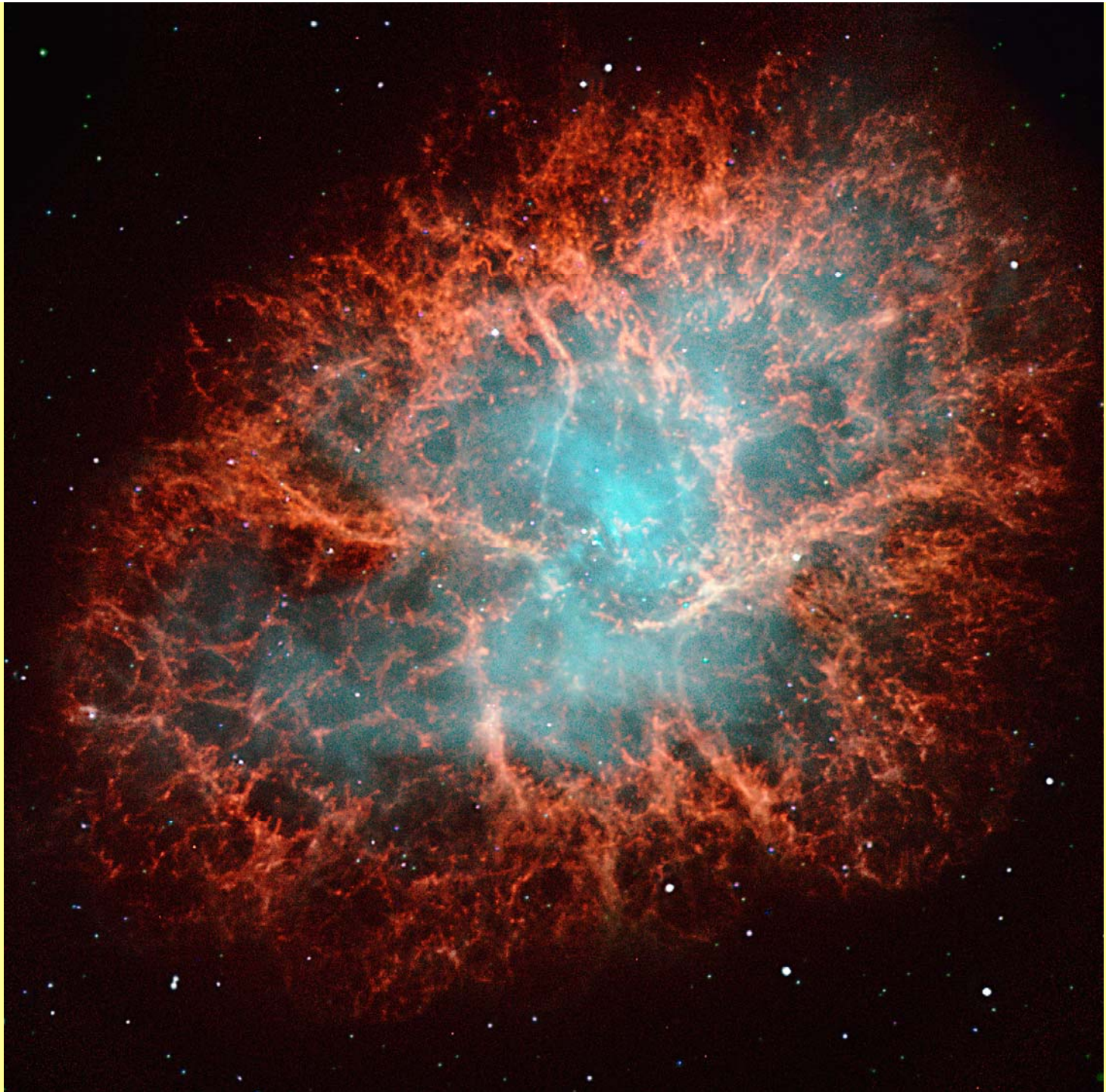


Composició (en massa)

Hidrogen	73.46 %
Heli	24.85 %
Oxigen	0.77 %
Carboni	0.29 %
Ferro	0.16 %
Sulfur	0.12 %
Neó	0.12 %
Nitrogen	0.09 %
Silici	0.07 %
Magnesi	0.05 %

Nebulosa
del Cranc

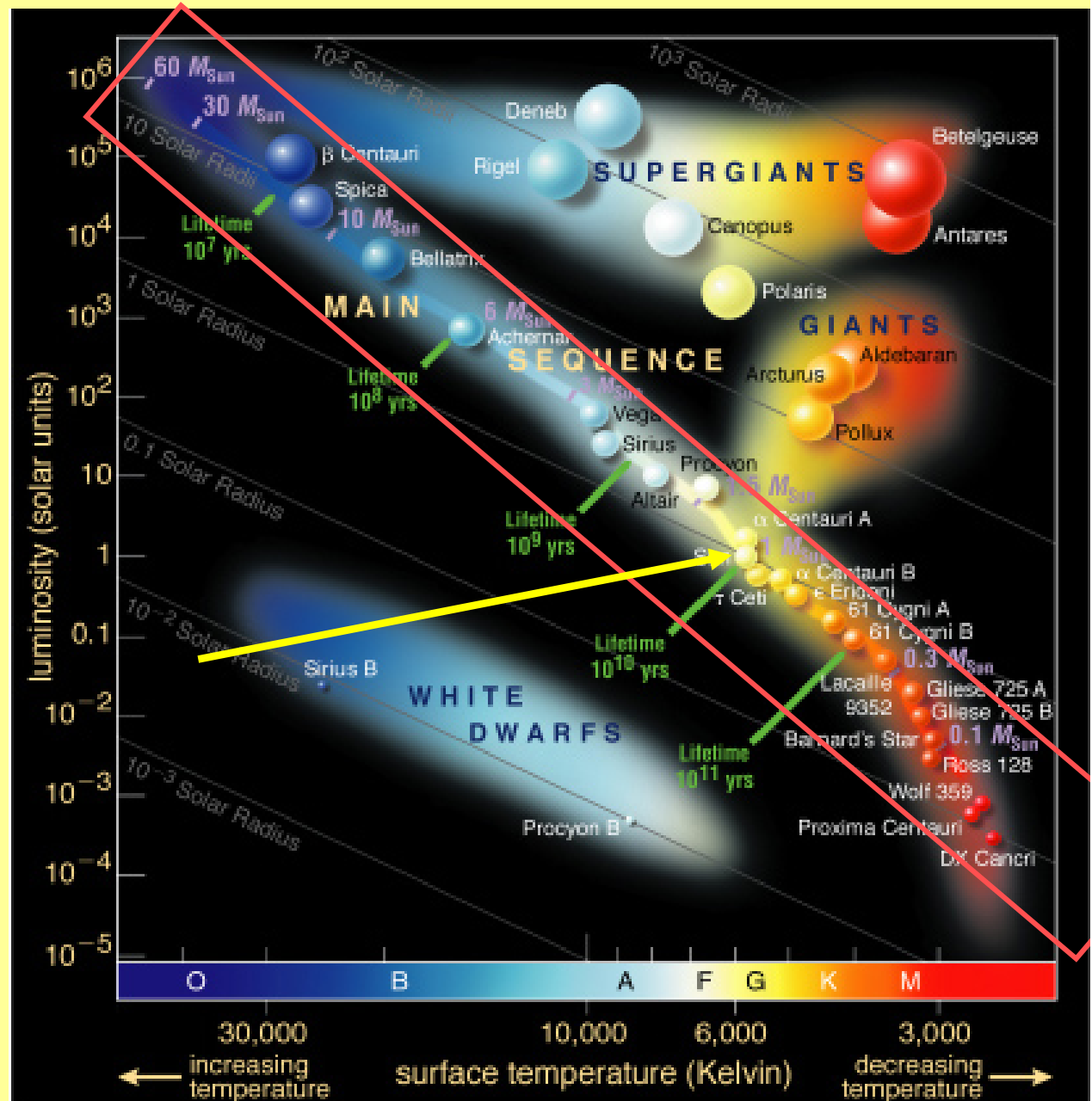
1054 d.C.

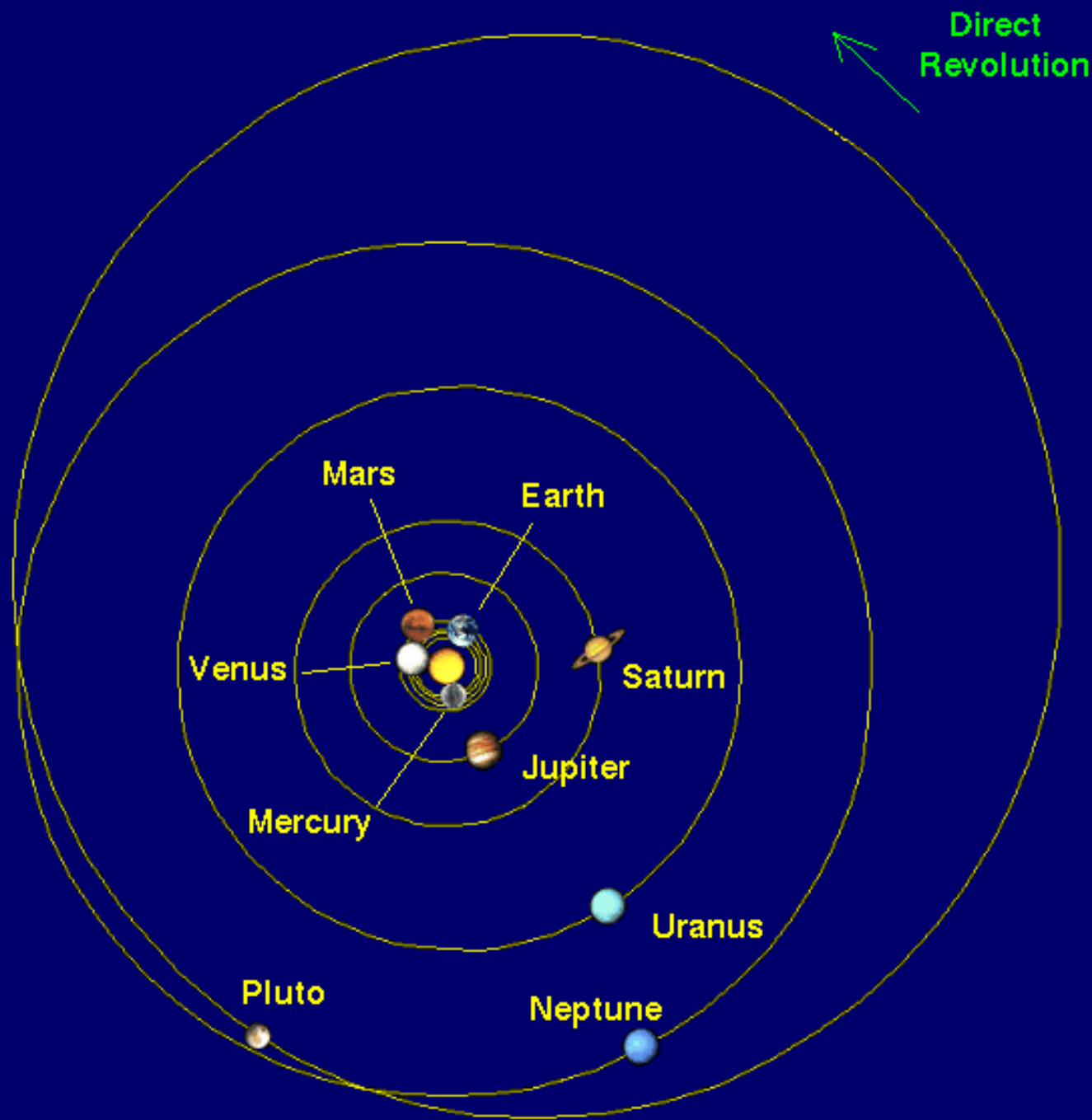


Actualment és a la seqüència principal: fusió hidrogen al nucli

5000 milions d'anys

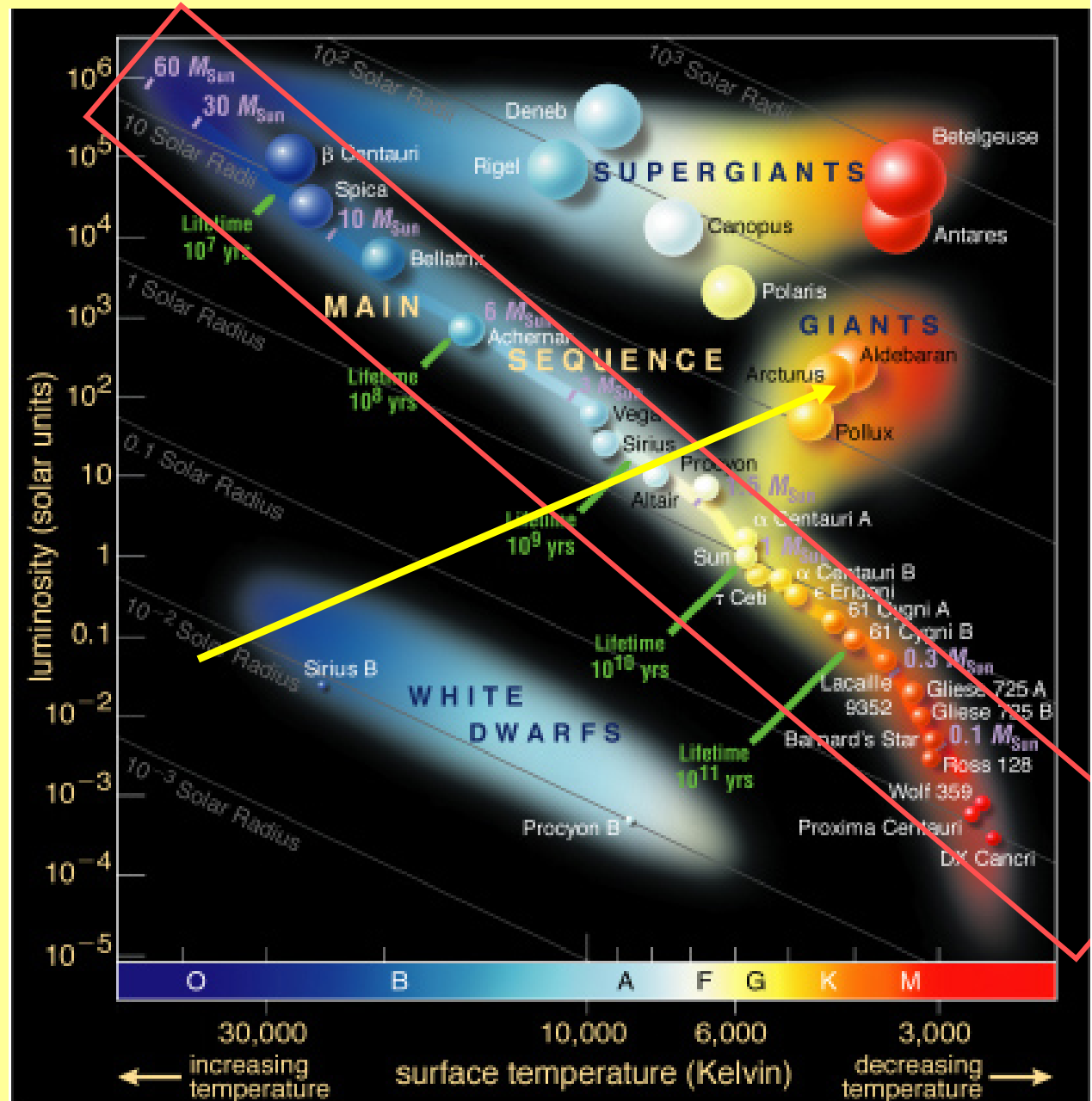
100 M_{Terra} \rightarrow energia





Actualment és a la seqüència principal: fusió hidrogen al nucli

+ 5000 milions d'anys



NGC 6543



Nana blanca de C+O, de la grandària de la Terra

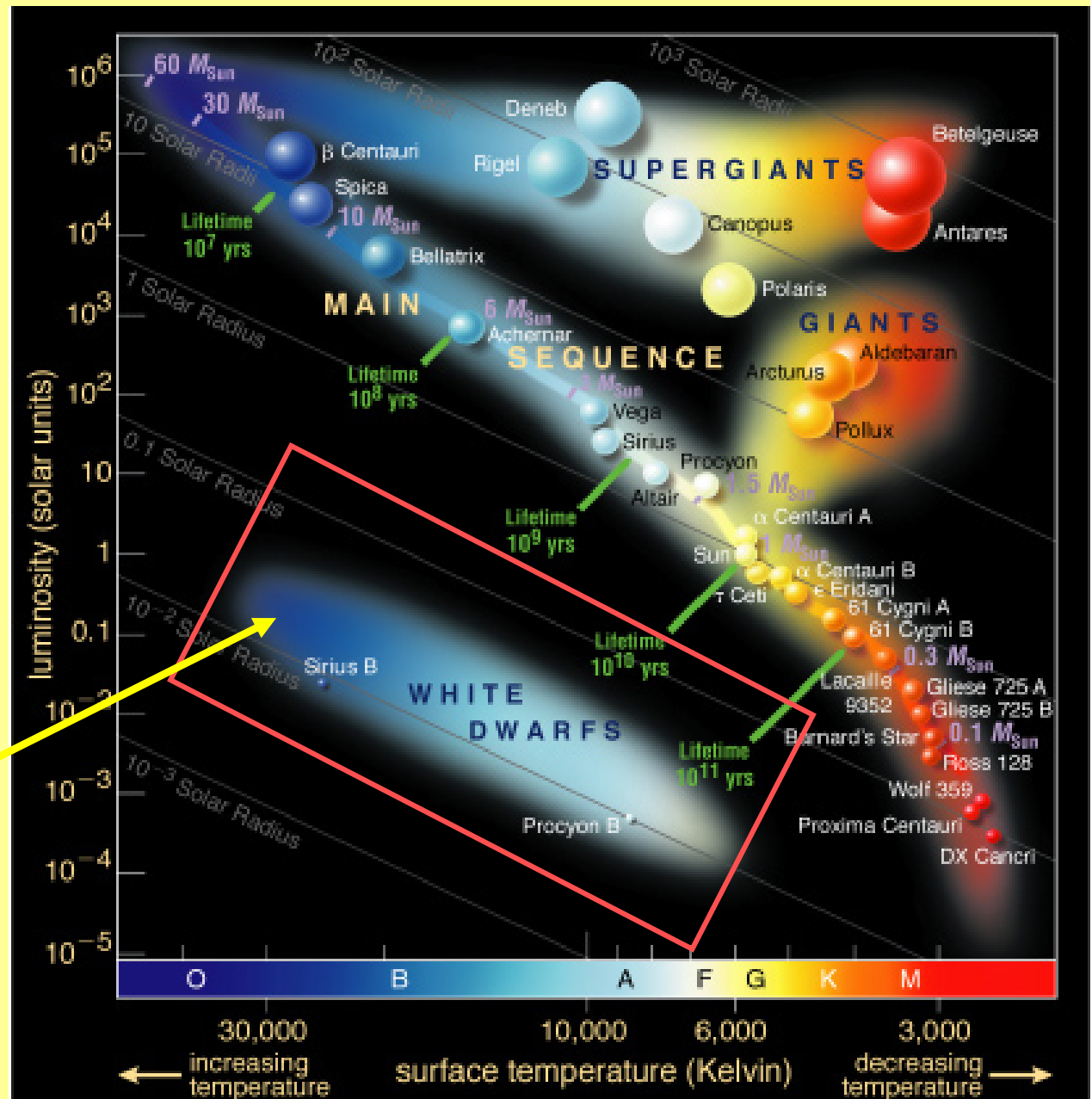
Nana blanca de C+O

Grandària de la Terra

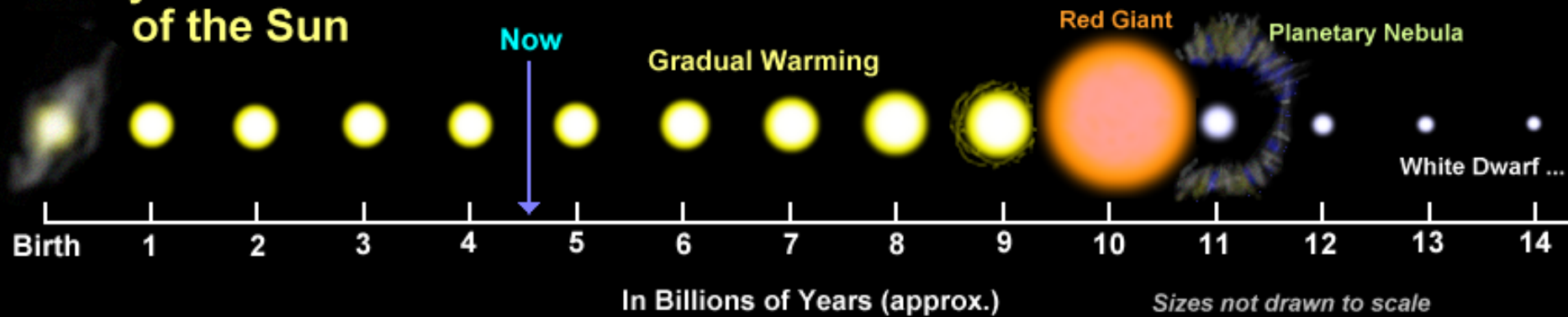
Molt densa

Es refredarà lentament

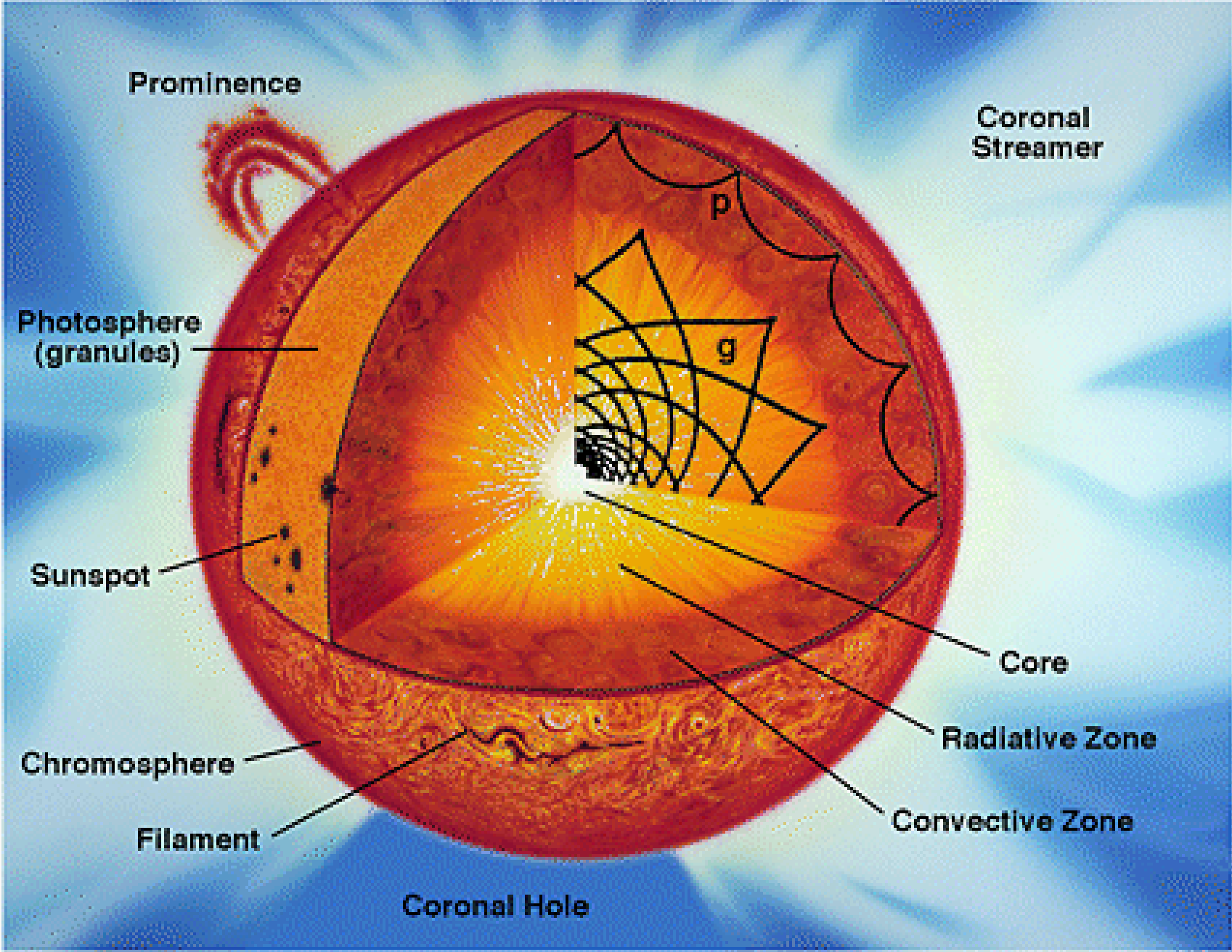
S'enfosquirà



Life Cycle of the Sun



Estructura



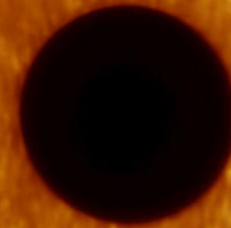
Period (min)

Fotosfera

enfosquiment
voves

granulació

UT 7 May 2003
1-m Solar Telescope on La Palma

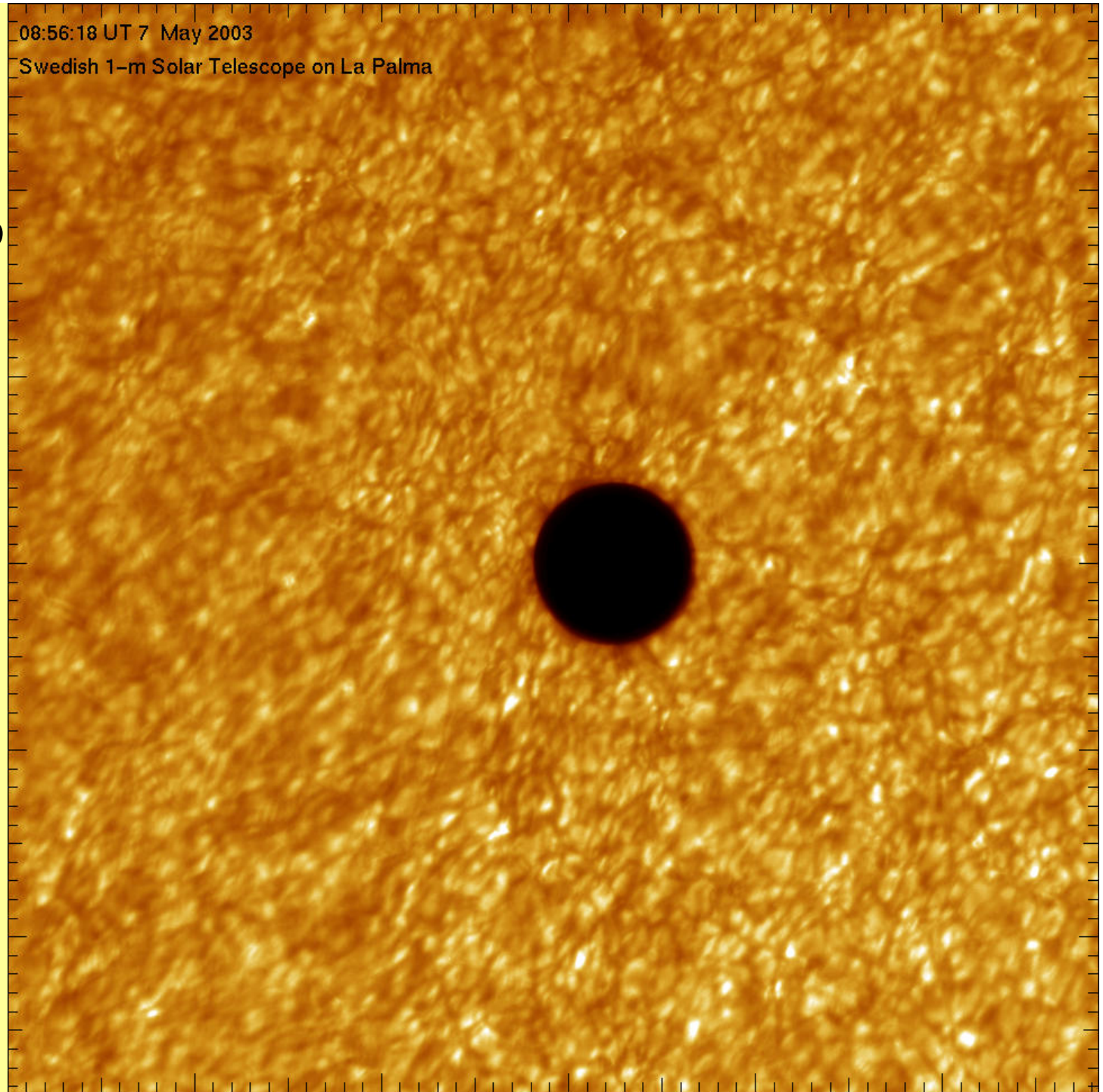


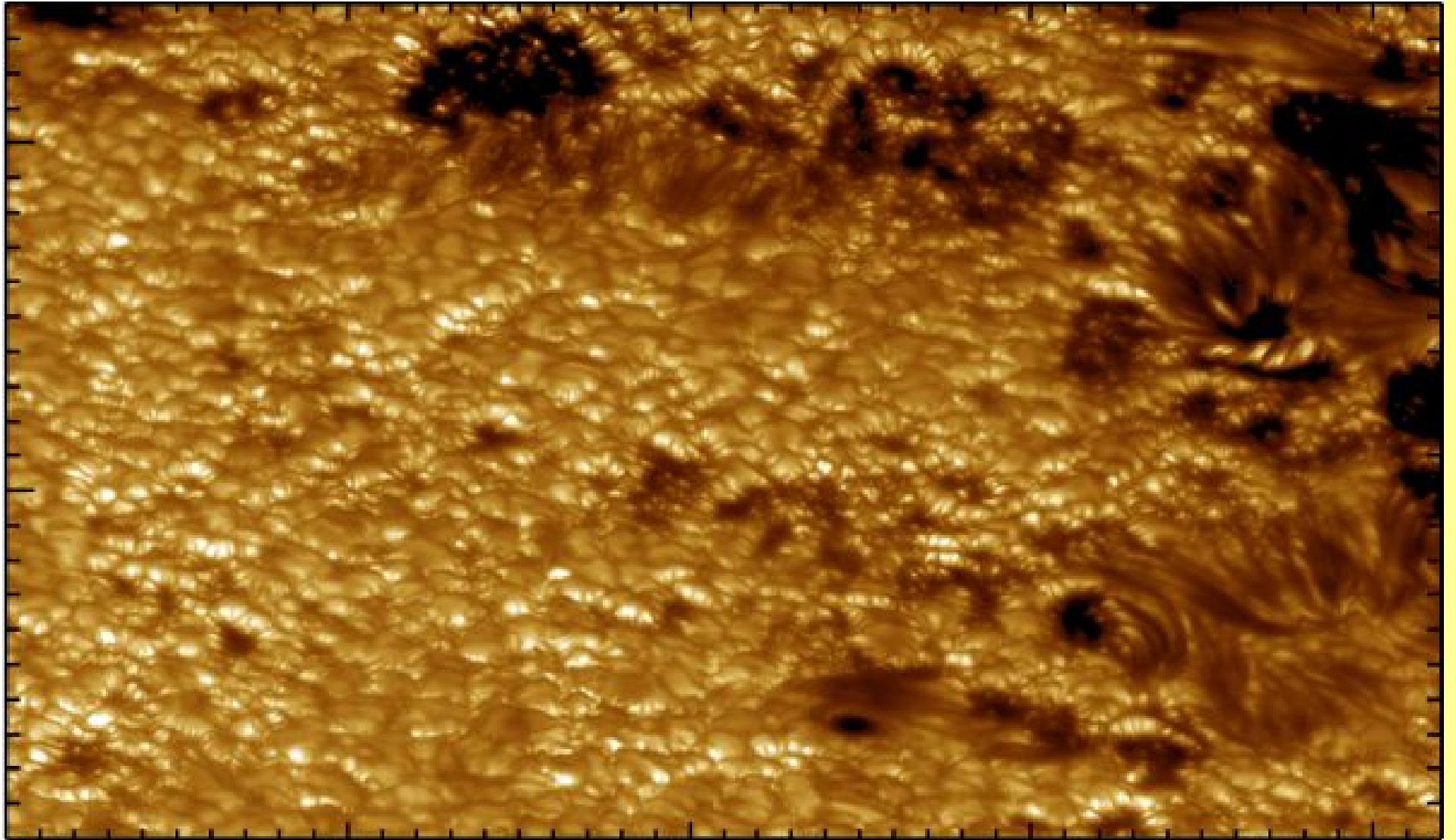
Fotosfera

granulació
supergranulació

08:56:18 UT 7 May 2003

Swedish 1-m Solar Telescope on La Palma







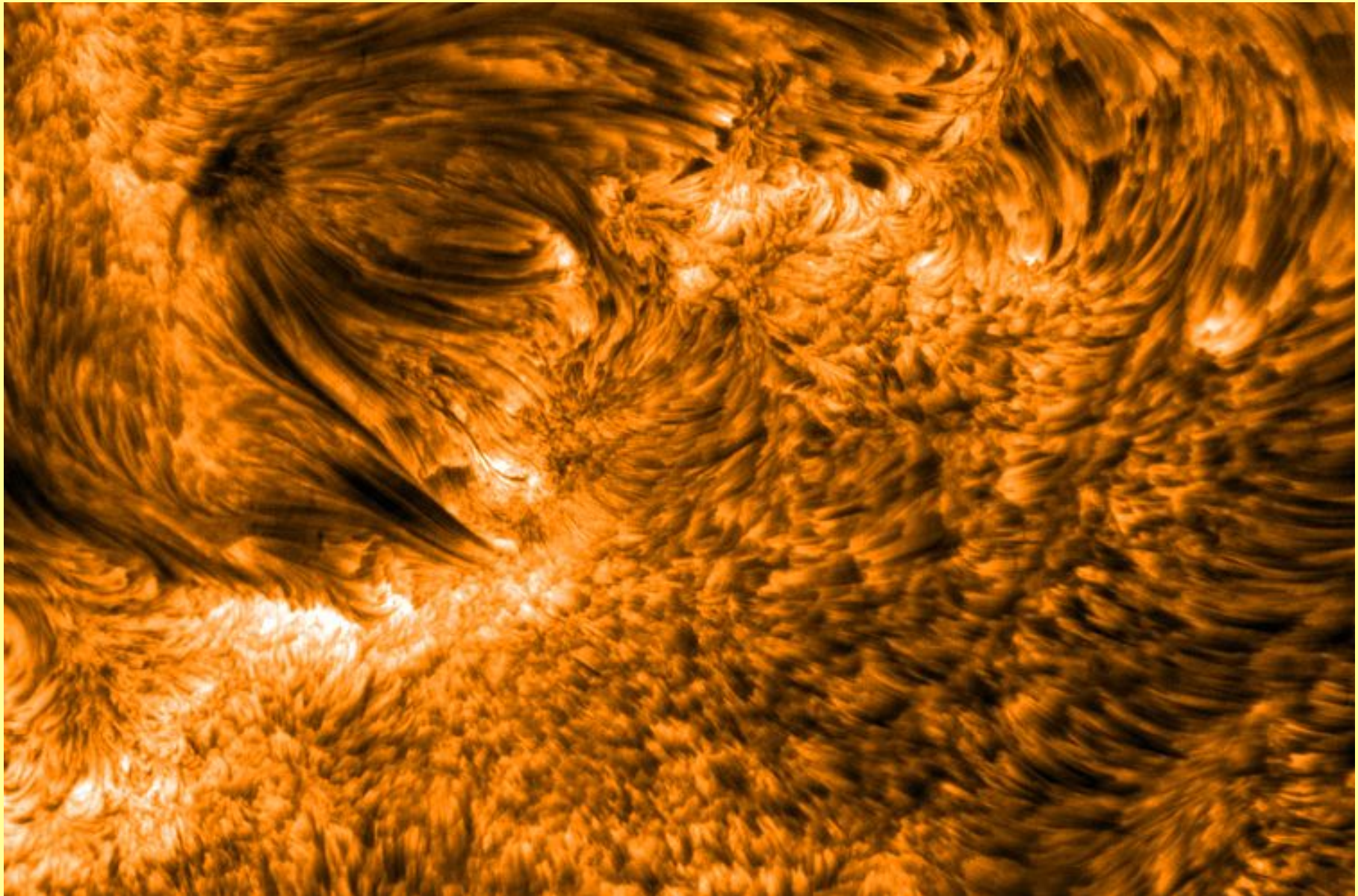
cromosfera



cromosfera

cromosfera

espículas



corona



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corona



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vent solar



Activitat

activitat solar

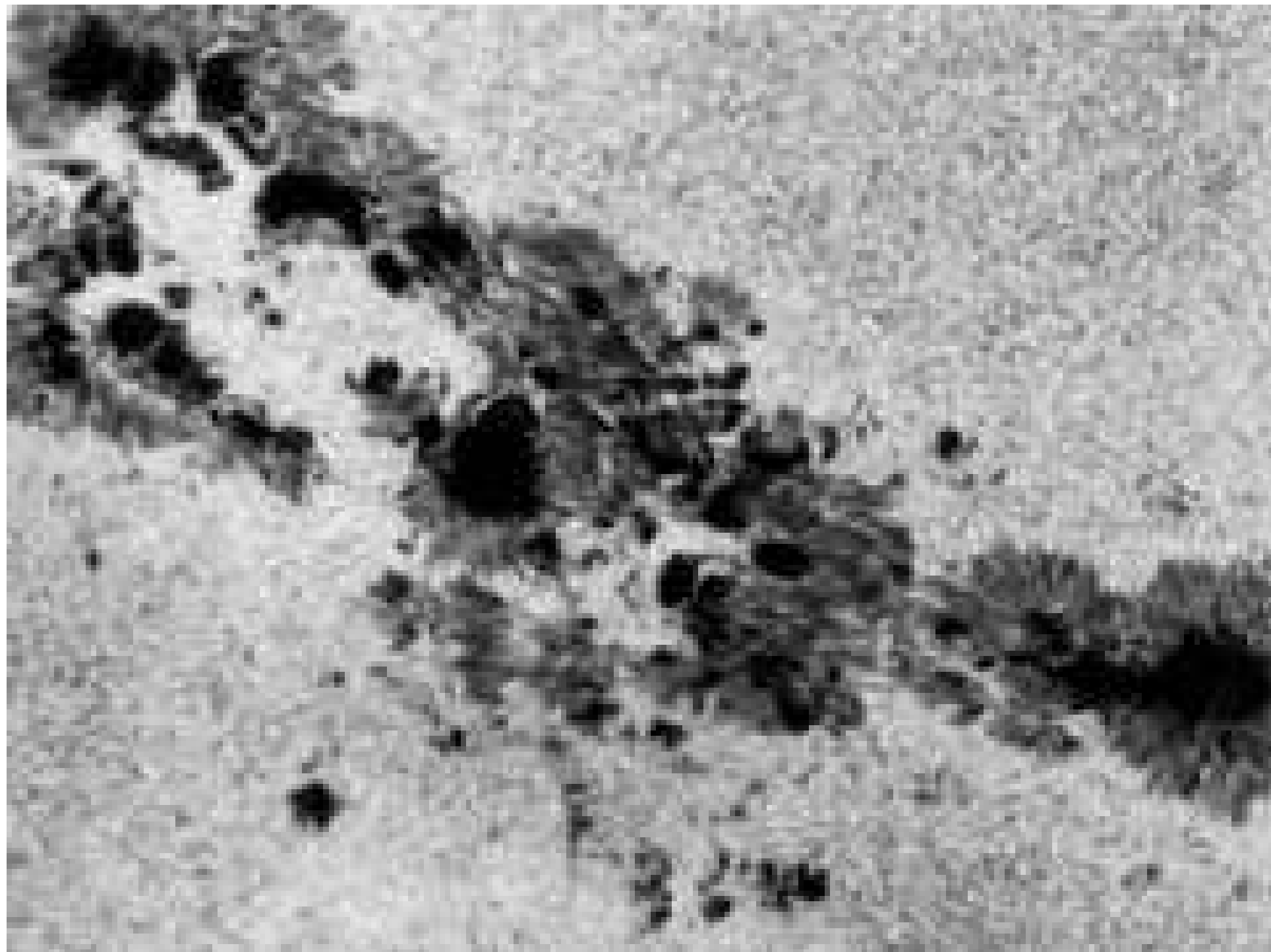
taques

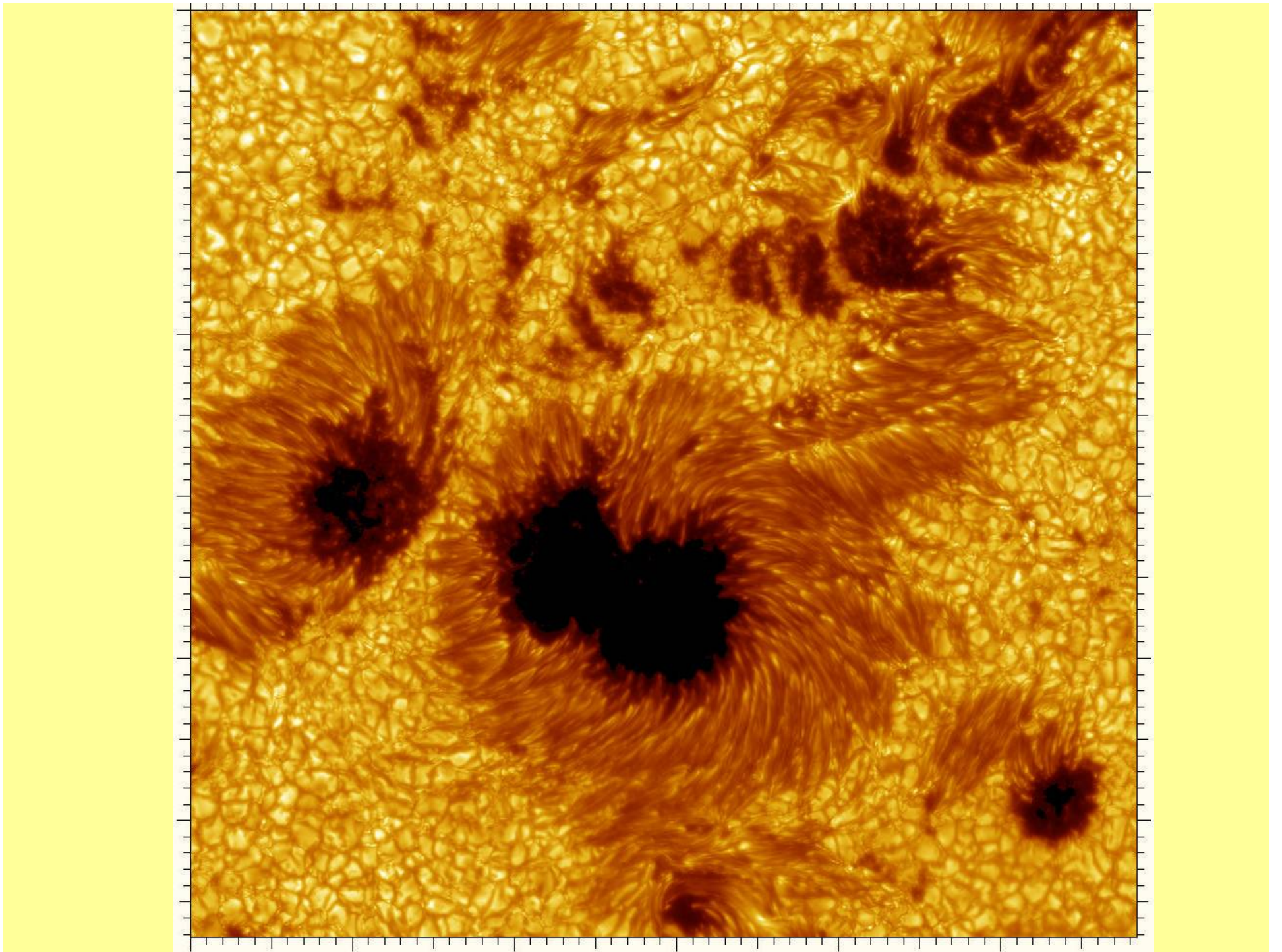


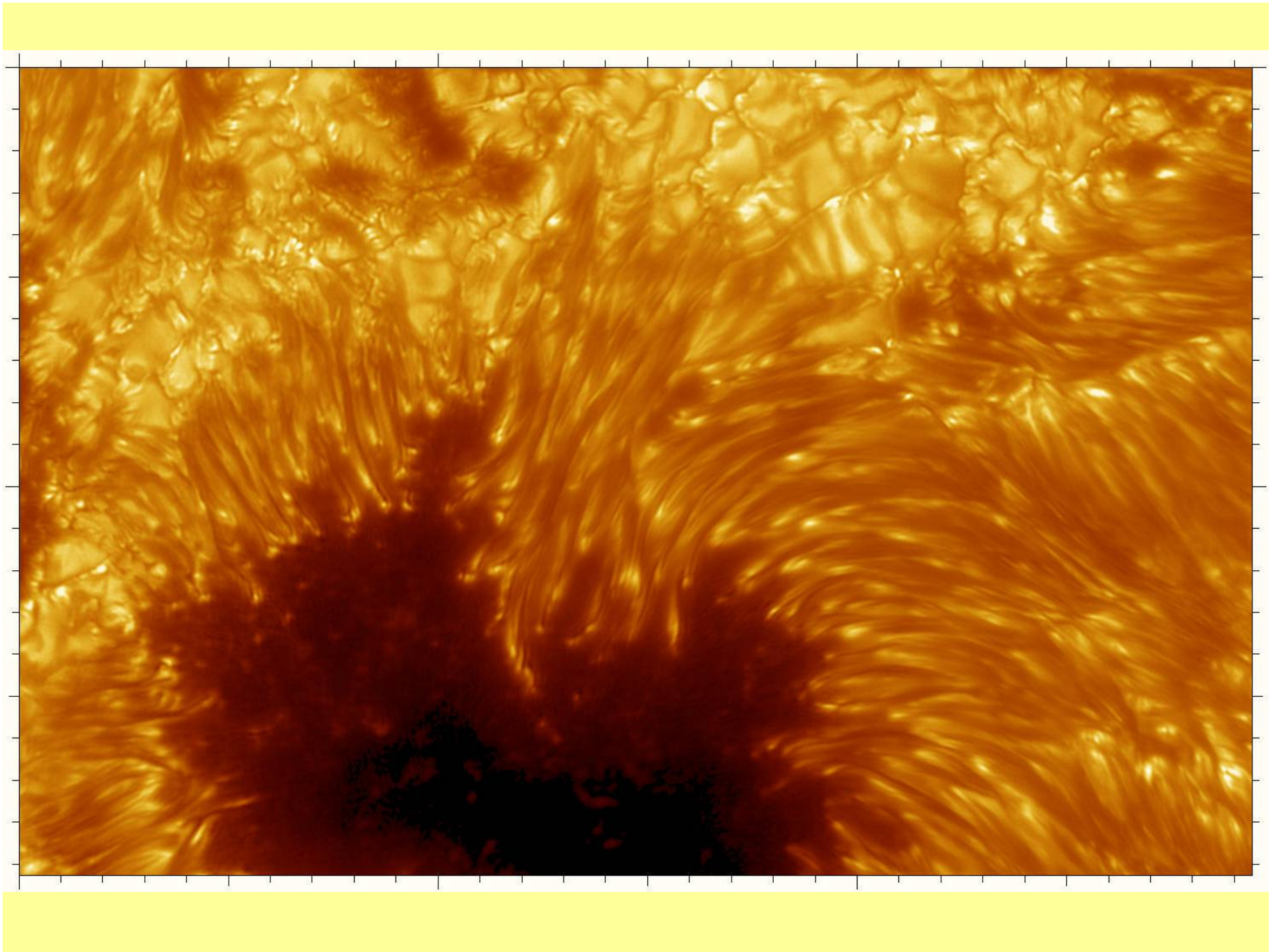


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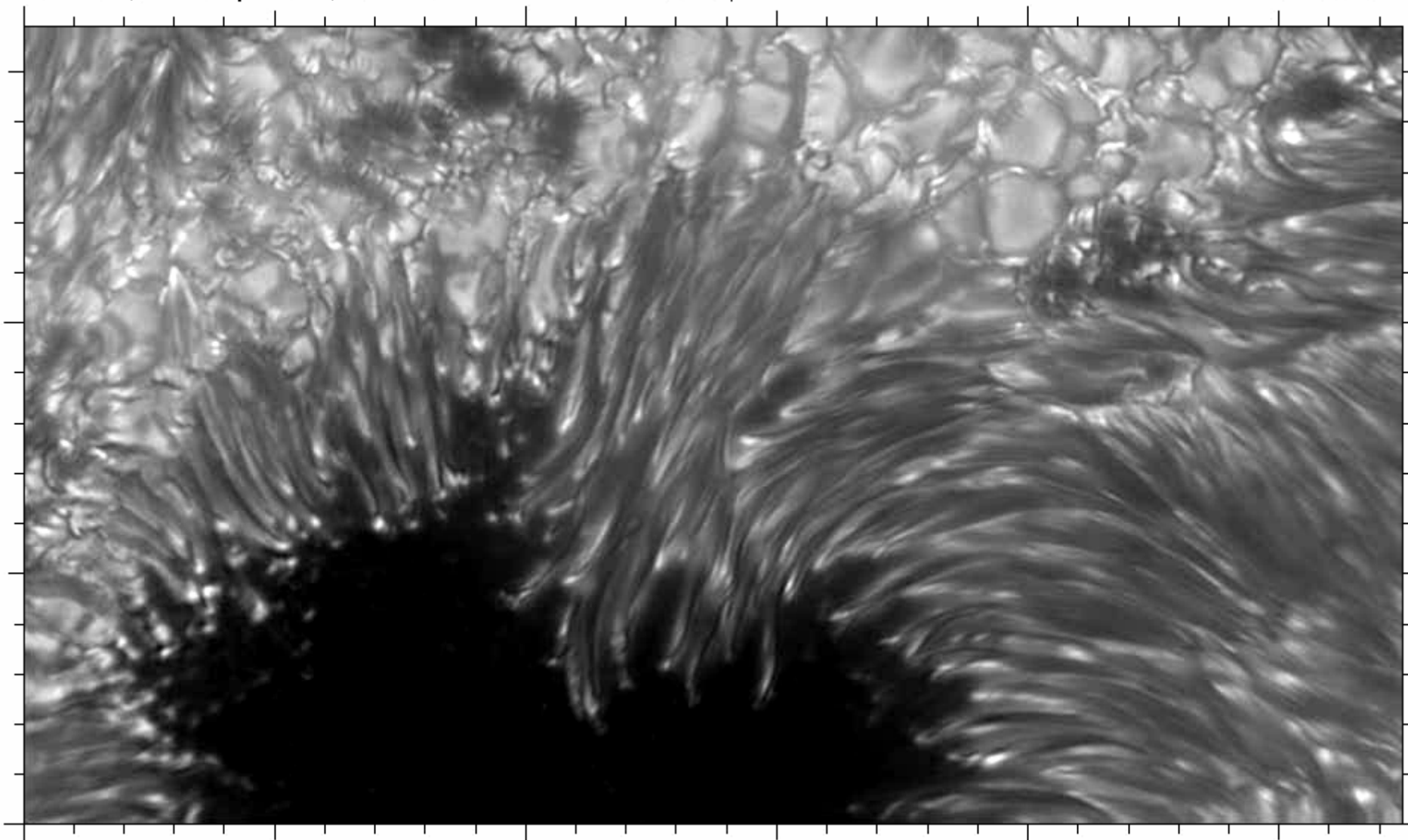






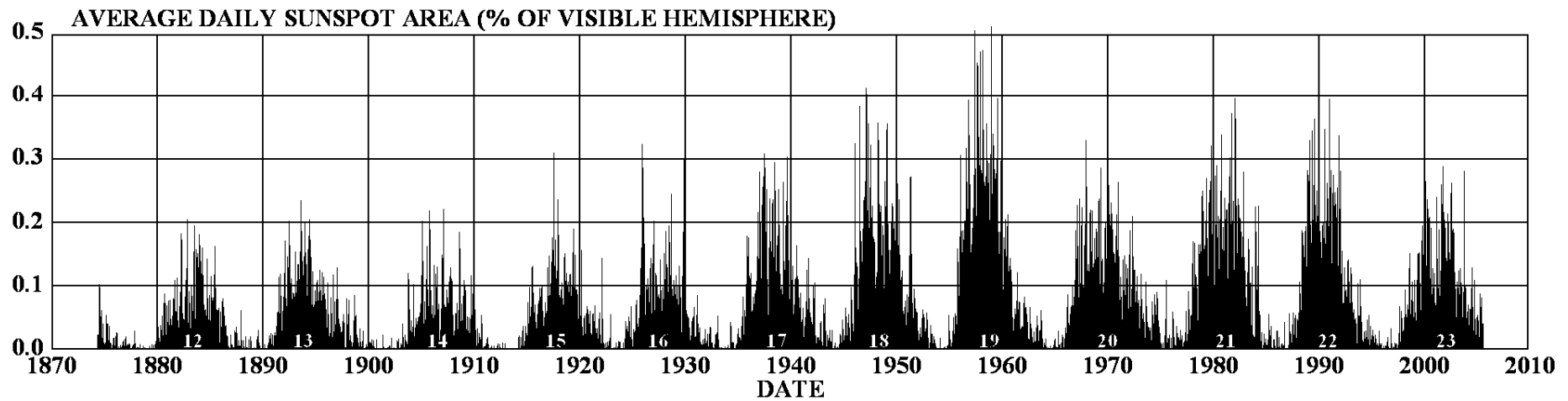
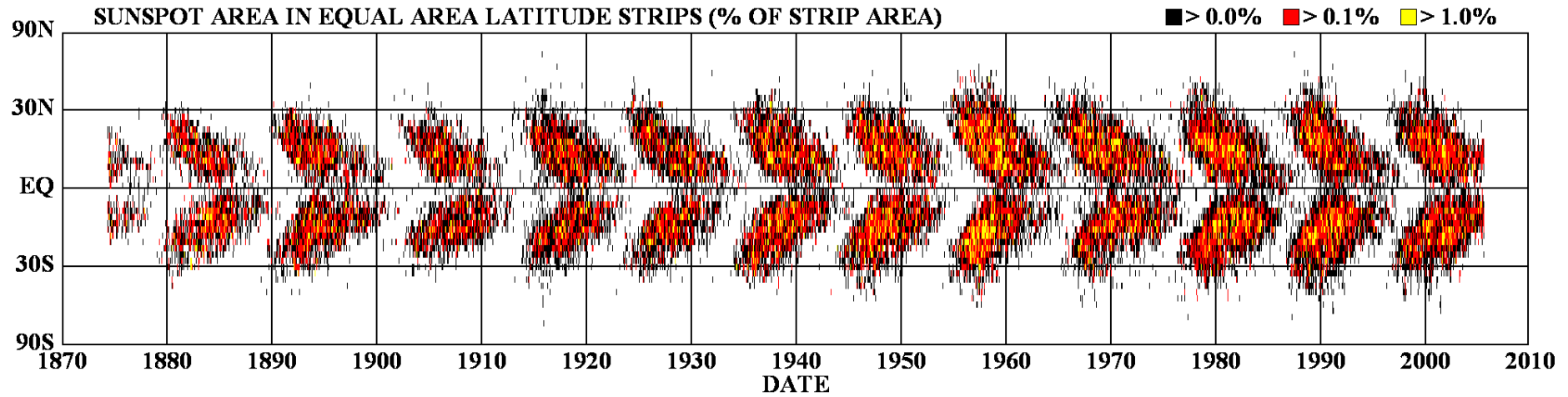
G-Band, 15 July 2002, Swedish 1-m Solar Telescope

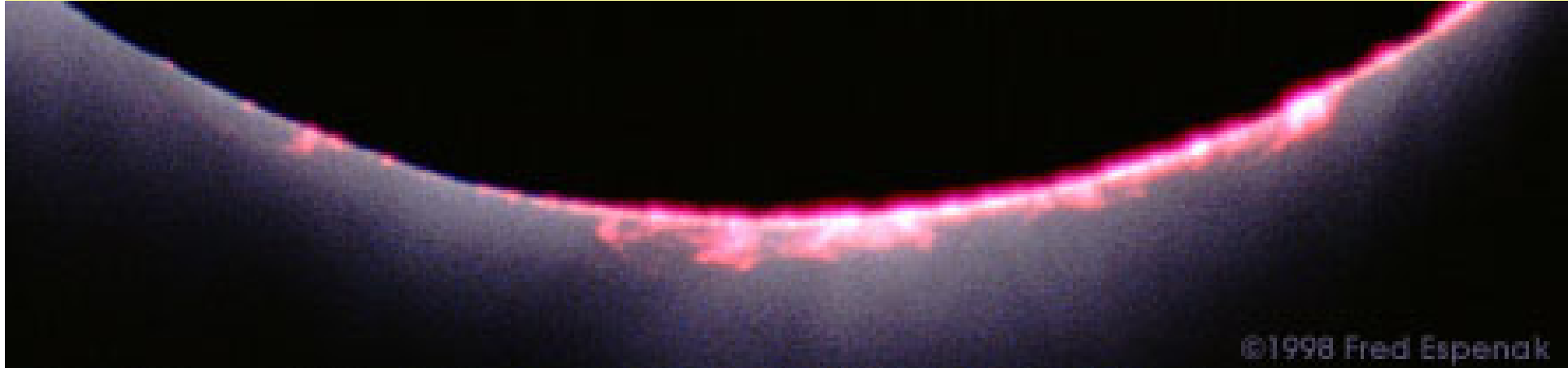
00:00:00



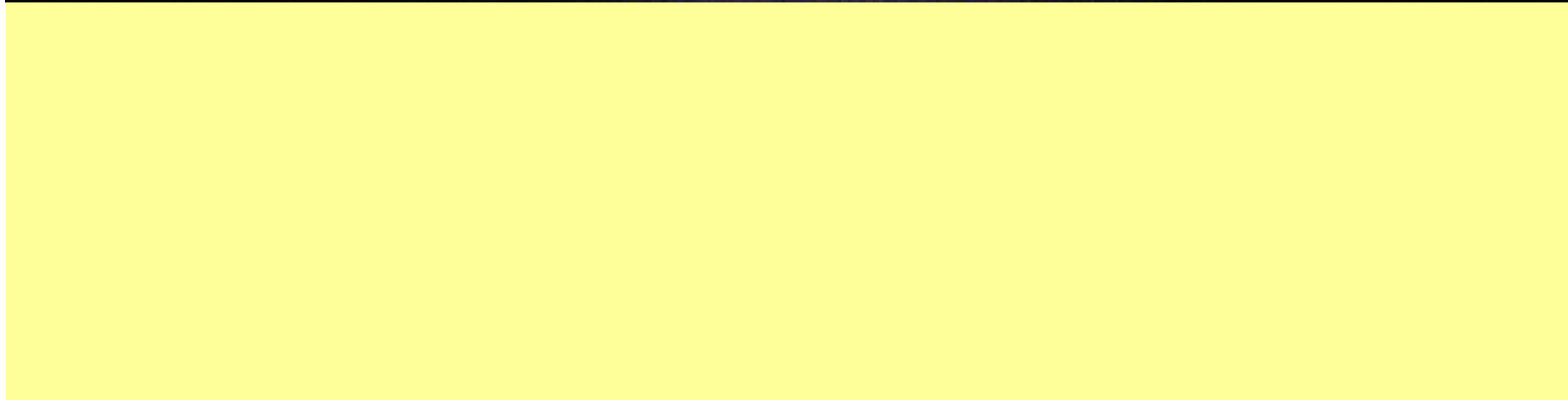
distance in units of 1000 kilometers

DAILY SUNSPOT AREA AVERAGED OVER INDIVIDUAL SOLAR ROTATIONS

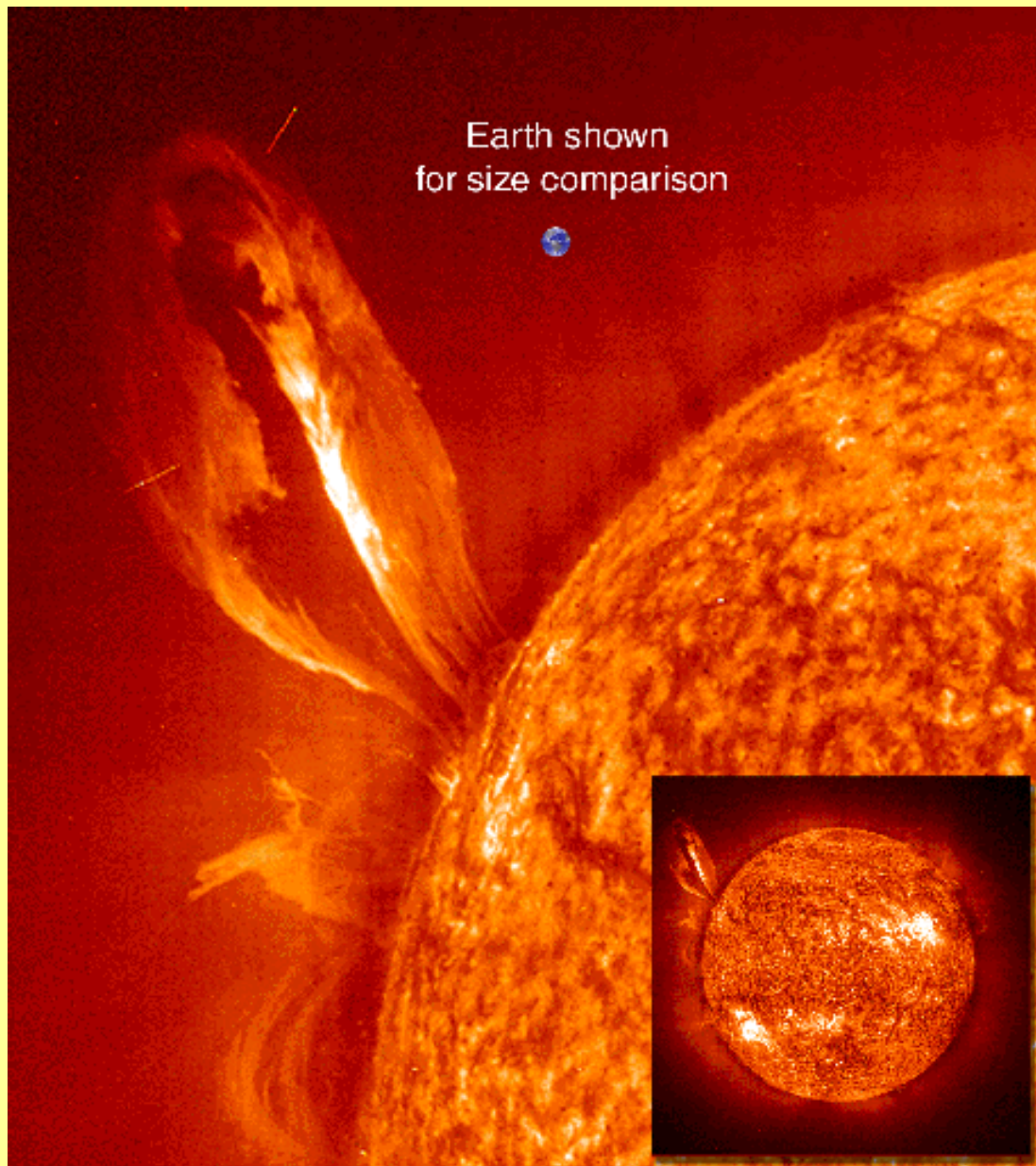


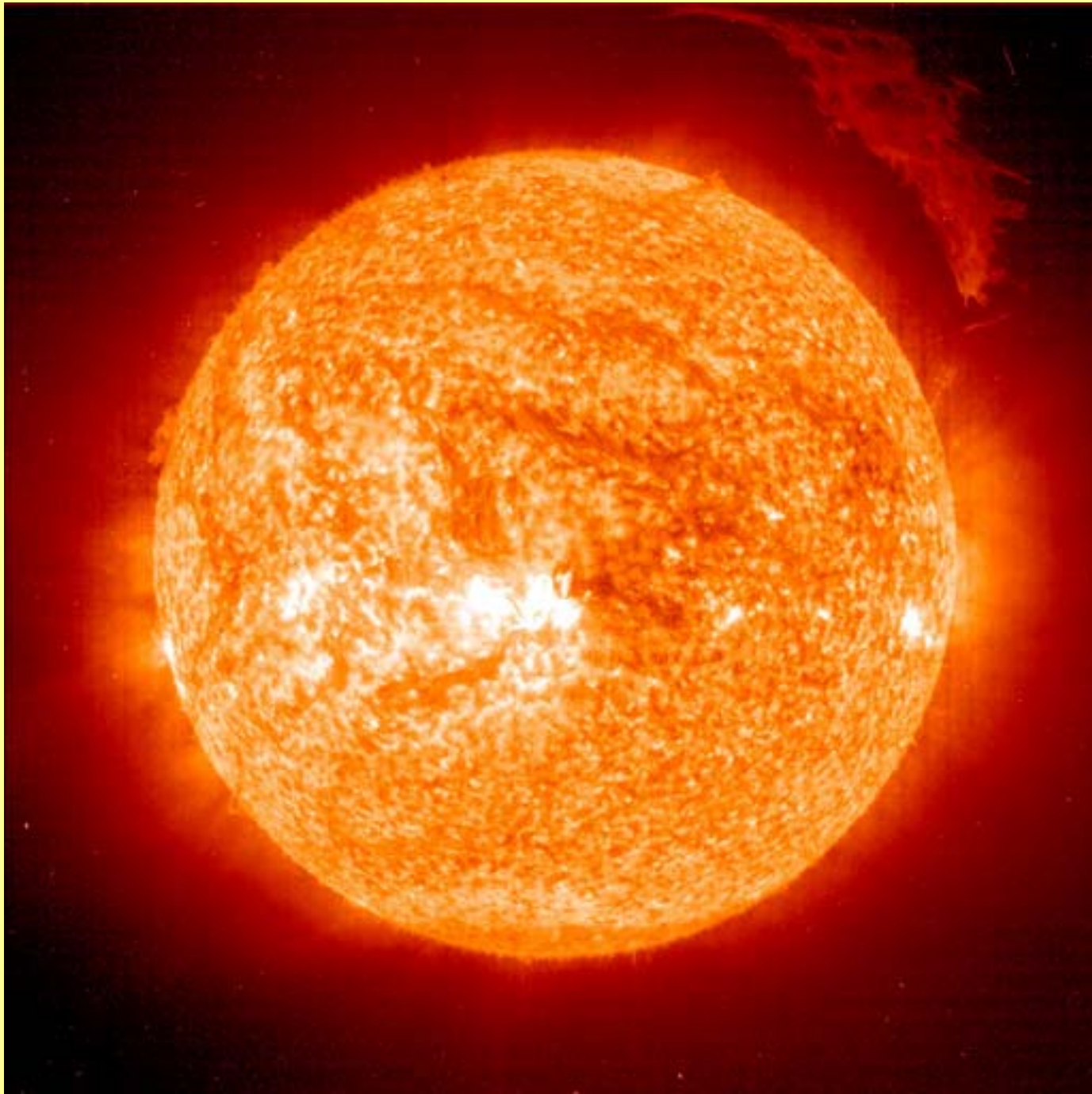


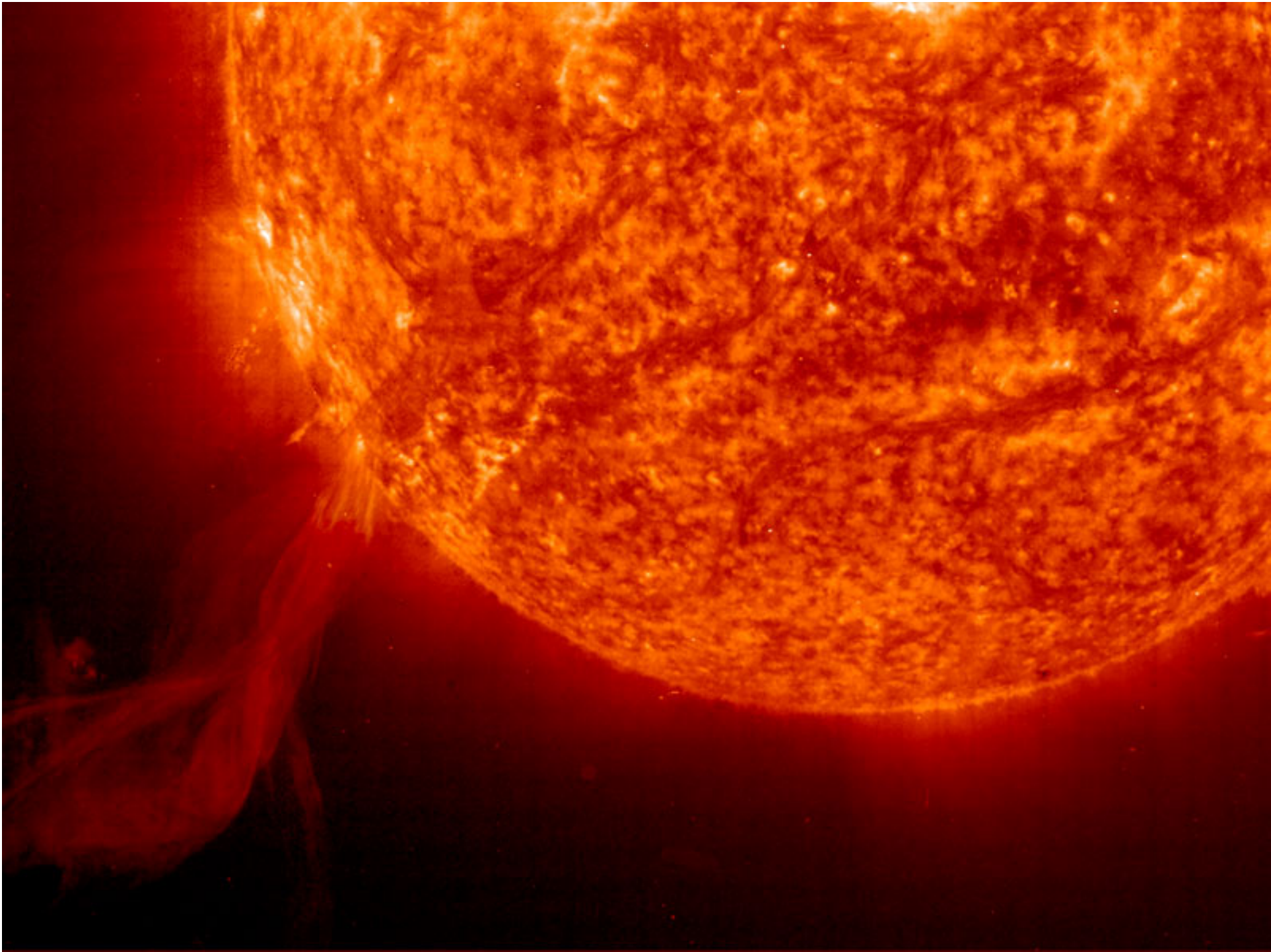
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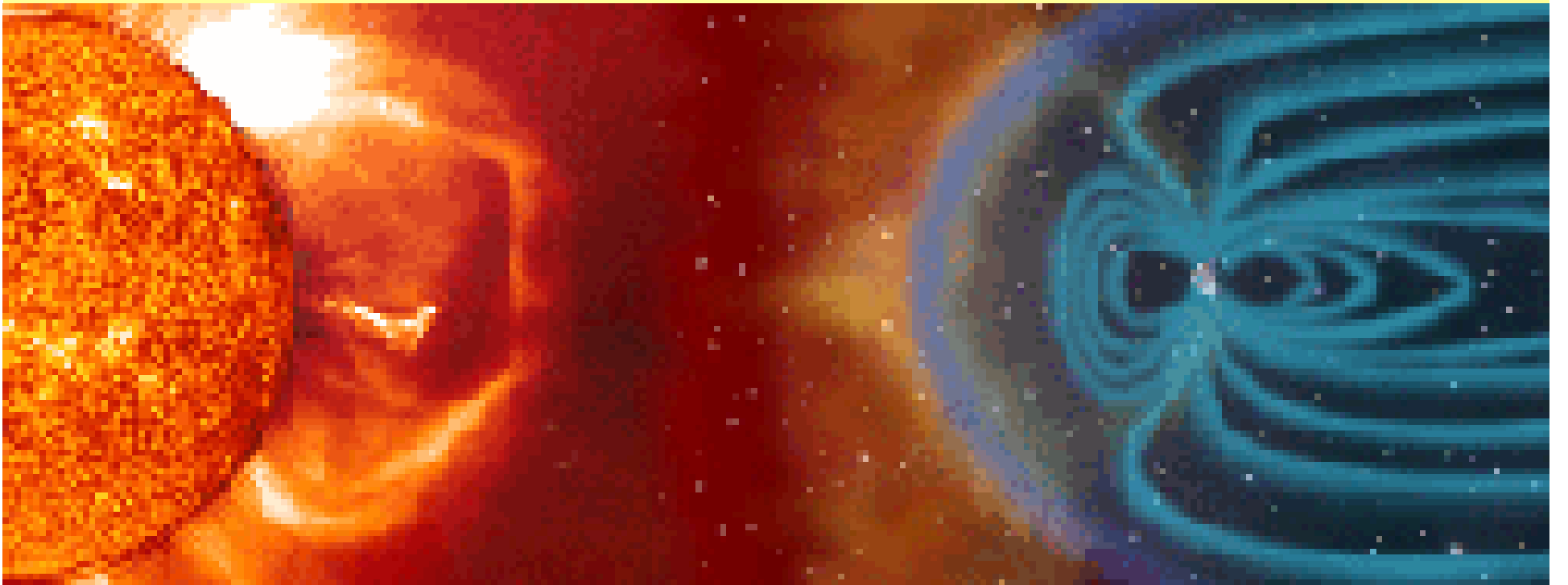


Earth shown
for size comparison



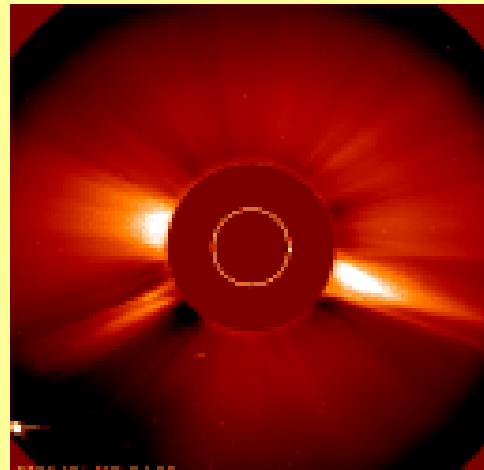
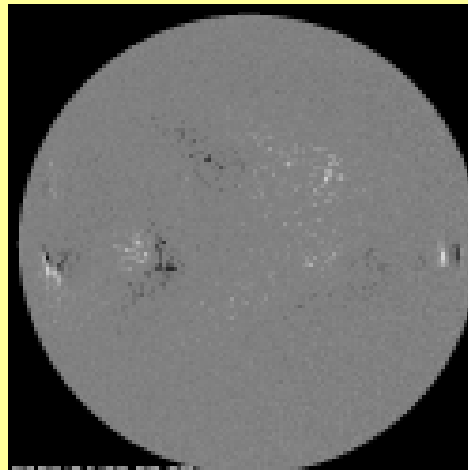
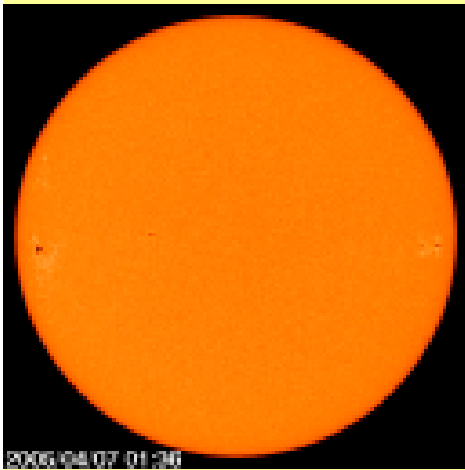
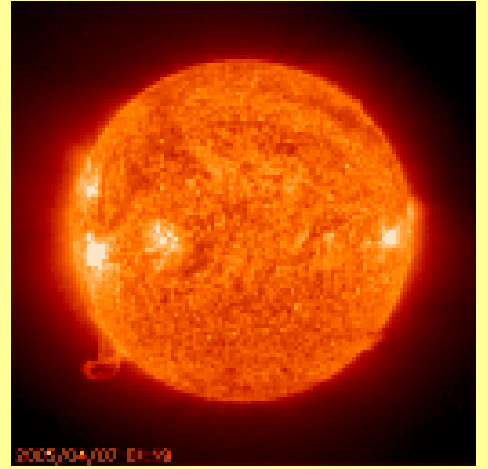
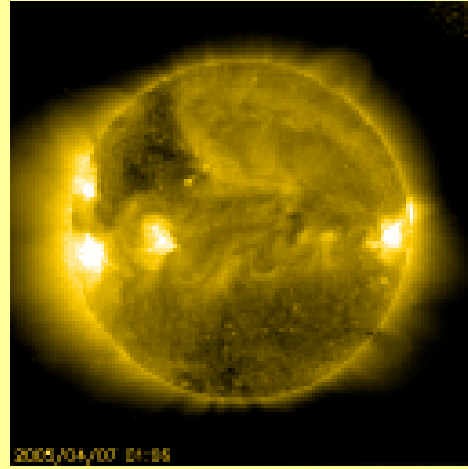
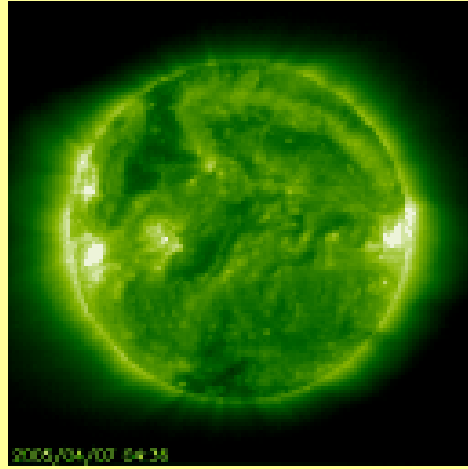
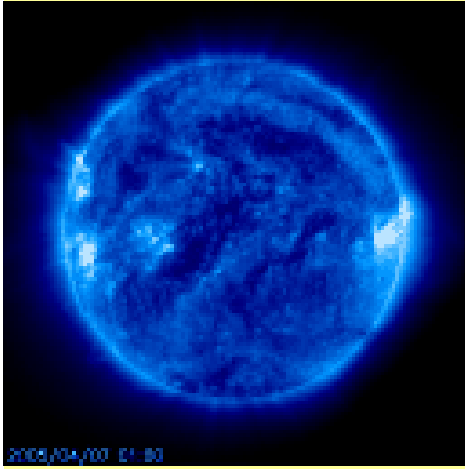


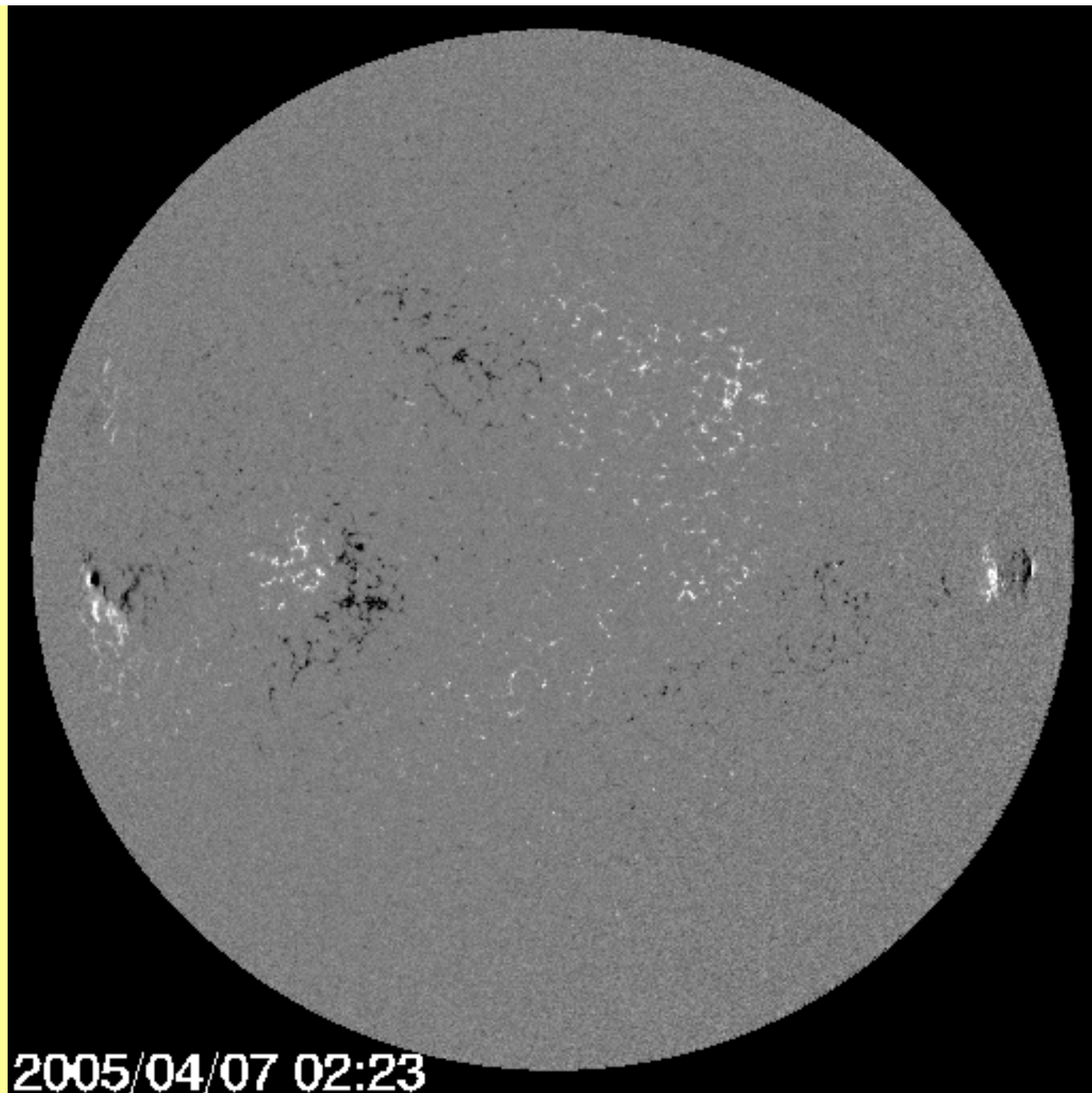




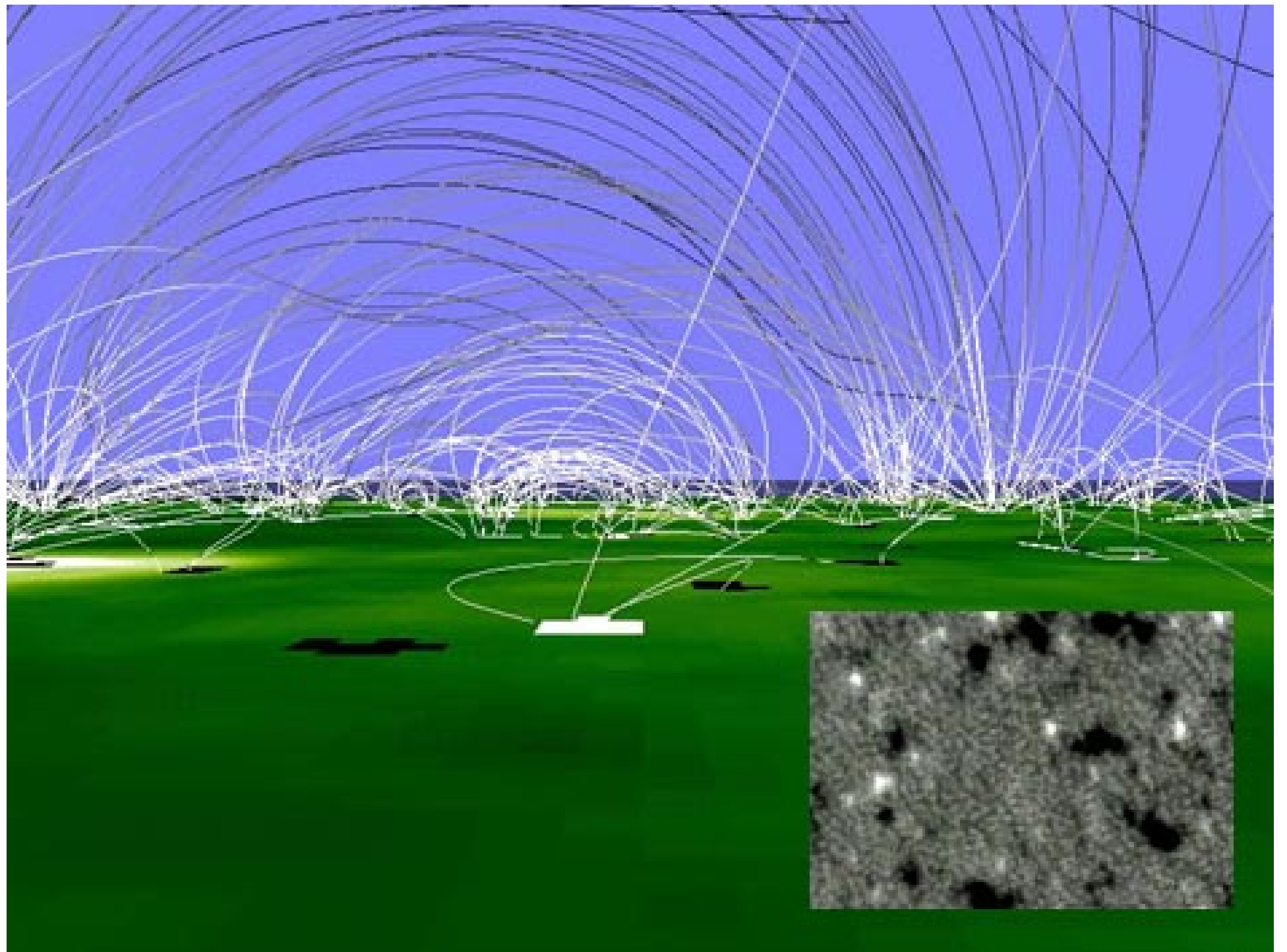








2005/04/07 02:23



Resum

- és una estrella 'grogua' (tipus G)
- és de segona-tercera generació
- és a la meitat de la seva vida
- morirà com una nana blanca (molt densa)
- nucli amb $H \rightarrow He$
- capa radiativa + capa convectiva
- atmosfera (fotosfera, cromosfera, corona)
- activitat magnètica (cicle 11 anys)
- taques, espícules, fàcules, prominències, ...

<http://www.michielb.nl/sun/kaft.htm>

Solar Physics: <http://science.nasa.gov/ssl/pad/solar/>

S.O.H.O.: <http://sohowww.estec.esa.nl/>

<http://seds.lpl.arizona.edu/nineplanets/nineplanets/sol.html>